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Author(s): Chung, Francis J.; Ola, Petri; Salo, Mikko; Tzou, Leo

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PARTIAL DATA INVERSE PROBLEMS FOR MAXWELL EQUATIONS VIA CARLEMAN ESTIMATES

FRANCIS J. CHUNG, PETRI OLA, MIKKO SALO, AND LEO TZOU

ABSTRACT. In this article we consider an inverse boundary value problem for the time-harmonic Maxwell equations. We show that the electromagnetic material parameters are determined by boundary measurements where part of the boundary data is measured on a possibly very small set. This is an extension of earlier scalar results of Bukhgeim-Uhlmann and Kenig-Sjöstrand-Uhlmann to the Maxwell system. The main contribution is to show that the Carleman estimate approach to scalar partial data inverse problems introduced in those works can be carried over to the Maxwell system.

RÉSUMÉ. Dans cet article nous considérons un problème inverse aux limites pour les équations de Maxwell harmoniques en temps. Nous montrons que les paramètres électromagnétiques sont déterminés par des mesures sur un très petit sous-ensemble du bord. Ces résultats pour le système de Maxwell sont une extension des résultats scalaires de Bukhgeim-Uhlmann et Kenig-Sjöstrand-Uhlmann. La contribution principale est de montrer que les méthodes d'estimations de Carleman de ces articles peuvent être généralisées au système de Maxwell.

1. INTRODUCTION

In this paper we discuss an inverse problem for the time-harmonic Maxwell equations with partial data, and show that the electromagnetic material parameters are determined by measurements on certain parts of the boundary. The result is new even for the case of bounded domains in \mathbb{R}^3 , but it will be convenient to formulate it more generally on compact manifolds with boundary.

Let (M, g) be a compact oriented Riemannian 3-manifold with C^∞ boundary. The electric and magnetic fields on M are described, respectively, by 1-forms E and H which satisfy the Maxwell equations in M :

$$(1.1) \quad \begin{cases} *dE = i\omega\mu H, \\ *dH = -i\omega\varepsilon E. \end{cases}$$

Here $\omega > 0$ is a fixed frequency, d is the exterior derivative, and $*$ is the Hodge star operator for the metric g . The material parameters ε and μ are assumed to be complex valued functions in $C^3(M)$ and to satisfy

$$\operatorname{Re}(\varepsilon) > 0, \operatorname{Re}(\mu) > 0.$$

The inverse problem is formulated in terms of partial measurements of the boundary tangential traces tE and tH of E and H . Here, the tangential trace of a k -form η is

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defined by

$$t : \eta \mapsto i^* \eta,$$

where $i : \partial M \rightarrow M$ is the inclusion map. If Γ_1 and Γ_2 are open subsets of ∂M , we define the partial Cauchy data set

$$C^{\Gamma_1, \Gamma_2} = \{(tE|_{\Gamma_1}, tH|_{\Gamma_2}); (E, H) \in H^2(M, \Lambda^1 M)^2 \text{ solves (1.1) and } \text{supp}(tE) \subset \Gamma_1\}.$$

As described in the end of Section 2, if ω is outside a discrete set of resonant frequencies, then the knowledge of C^{Γ_1, Γ_2} is equivalent with knowing the partial admittance map

$$\Lambda^{\Gamma_1, \Gamma_2} : f \mapsto tH|_{\Gamma_2},$$

where f is a 1-form on ∂M with $\text{supp}(f) \subset \Gamma_1$, and (E, H) is the unique solution of (1.1) satisfying $tE = f$ on ∂M . This corresponds to prescribing tE on Γ_1 and measuring tH on Γ_2 . The inverse problem is to determine the coefficients ε and μ from the knowledge of C^{Γ_1, Γ_2} for some choices of Γ_j .

One may think of the above inverse problem for Maxwell equations as a generalization to systems of the inverse conductivity problem introduced by Calderón [Ca80]. In the case where M is a bounded domain in \mathbb{R}^3 and g is the Euclidean metric, it was proved in [OPS93] that measurements on the full boundary ($\Gamma_1 = \Gamma_2 = \partial M$) determine ε and μ uniquely. Earlier results include [SIC92, SU92, CP92], and a simplified proof was presented in [OS96]. Stability results for this inverse problem are in [Ca10, Ca11], boundary determination results are in [Mc97, JM00], the case of chiral media is considered in [Mc00], and a recent uniqueness result for C^1 coefficients is given in [CZ13].

The inverse problem for Maxwell equations on manifolds was discussed in [OPS03]. The first uniqueness results in the non-Euclidean case were given in [KSU09], again for the full data case $\Gamma_1 = \Gamma_2 = \partial M$, when M is an admissible manifold in the following sense.

Definition 1.1. *A compact Riemannian manifold (M, g) with smooth boundary is called admissible if (M, g) is embedded in (T, g) where $T = \mathbb{R} \times M_0$ and $g = c(e \oplus g_0)$, where c is a smooth positive function, (\mathbb{R}, e) is the Euclidean line, and (M_0, g_0) is a simple manifold (a compact manifold with smooth strictly convex boundary such that the exponential map \exp_p is a diffeomorphism onto M_0 for each $p \in M_0$).*

Locally, admissibility implies that in some local coordinates $x = (x_1, x')$ the metric takes the form

$$g(x_1, x') = c(x_1, x') \begin{pmatrix} 1 & 0 \\ 0 & g_0(x') \end{pmatrix}$$

where $g_0(x')$ is some positive definite $(n-1) \times (n-1)$ matrix. Conversely, any metric that takes this form in some local coordinates is admissible provided that g_0 is simple. Admissible manifolds include compact submanifolds of Euclidean space, hyperbolic space, and S^3 minus a point. They also include sufficiently small submanifolds of conformally flat manifolds and warped products. For more information see [DKSU09, AFGR16].

In this article we will improve the results of [OPS93, KSU09] by considering the partial data problem where $\Gamma_1 = \partial M$ but Γ_2 is an appropriately chosen open subset of ∂M which is defined by a limiting Carleman weight (LCW) of the manifold M . The notion of LCWs was introduced in connection with the Calderón problem with partial data in [KSU07] in Euclidean space. LCWs were analysed in detail in [DKSU09] also on manifolds. In particular, the class of admissible manifolds emerged in [DKSU09] as a natural class where LCWs exist and one may expect to be able to solve related inverse problems. Note

that any admissible manifold has global coordinates (x_1, x') where x_1 is the Euclidean direction. It was shown in [DKSU09] that $\varphi(x) = x_1$ is a natural LCW in this setting, and in this paper we will always assume that the LCW is given by $\varphi(x) = x_1$.

We define the “front face” of the boundary with respect to the LCW φ by

$$F_\varphi := \{p \in \partial M \mid \langle d\varphi(p), \nu \rangle \geq 0\}.$$

Here ν is the 1-form corresponding to the unit outer normal of $\partial\Omega$. Our main result is the following theorem.

Theorem 1.1. Let (M, g) be an admissible 3-manifold and $\varphi(x) = x_1$ a LCW on M . Let $\varepsilon_j, \mu_j \in C^3(M)$ be complex functions with positive real part, and let $\omega > 0$. Suppose that the corresponding Cauchy data sets satisfy

$$C_1^{\partial M, \tilde{F}} = C_2^{\partial M, \tilde{F}}$$

for some open neighborhood \tilde{F} of F_φ in ∂M . Assume in addition that $\varepsilon_1 = \varepsilon_2$ and $\mu_1 = \mu_2$ to second order on ∂M . Then $\varepsilon_1 \equiv \varepsilon_2$ and $\mu_1 \equiv \mu_2$ in M .

We mention a particular case of the above theorem in \mathbb{R}^3 ($\text{ch}(\bar{\Omega})$ denotes convex hull).

Theorem 1.2. Let $\Omega \subset \mathbb{R}^3$ be a bounded domain with C^∞ boundary, let $\varepsilon_j, \mu_j \in C^3(\bar{\Omega})$ be complex functions with positive real part, and let $\omega > 0$. Given $x_0 \in \mathbb{R}^3 \setminus \text{ch}(\bar{\Omega})$, define

$$F(x_0) = \{x \in \partial\Omega; (x - x_0) \cdot \nu(x) \leq 0\}.$$

Suppose that the corresponding Cauchy data sets satisfy

$$C_1^{\partial\Omega, \tilde{F}} = C_2^{\partial\Omega, \tilde{F}}$$

for some open neighborhood \tilde{F} of $F(x_0)$ in $\partial\Omega$. Assume in addition that $\varepsilon_1 = \varepsilon_2$ and $\mu_1 = \mu_2$ to second order on $\partial\Omega$. Then $\varepsilon_1 \equiv \varepsilon_2$ and $\mu_1 \equiv \mu_2$ in Ω .

This result involves the logarithmic Carleman weight in \mathbb{R}^3 and corresponds to one of the results of [KSU07] for the partial data problem for the Schrödinger equation. Note that if Ω is strictly convex, then any open subset of $\partial\Omega$ can serve as \tilde{F} above.

A standard method to study such partial data problems is to apply the idea developed in [BU02, KSU07] for scalar equations, where Carleman estimates with suitable positive boundary terms were used to suppress the unknown information. The partial data inverse problem for Maxwell equations presents several challenges when one tries to apply these methods directly. The first difference is the fact that the principal part of this equation is a first order system rather than a second order scalar equation, and Carleman estimates with boundary terms for systems seem to be more involved than those for scalar equations (see [El08, ST09] for some such estimates). Secondly, the fact that ellipticity is somewhat “hidden” in the system makes it difficult to obtain an estimate similar to the one used in [KSU07] which could be directly applied to the Maxwell system.

We circumvent these difficulties by first using ideas from [OS96, OPS03, KSU09] to show that the problem reduces to constructing suitable solutions to a Hodge Dirac system

$$(P + W)u = 0, \quad tu|_{\partial M \setminus \tilde{F}} = 0.$$

We will then apply the recent work [CST17] where the authors studied Carleman estimates and complex geometrical optics solutions for the Hodge Laplace operator with relative and absolute boundary conditions. The crucial point is that the Carleman estimate with

boundary terms proved in [CST17] for the Hodge Laplacian is sufficiently powerful to yield information also in the Maxwell case. However, we stress that the present situation for Maxwell equations does not reduce to that of [CST17], since the boundary measurements for Maxwell determine in a sense only part of the boundary measurements for the Hodge system. It is essential to use the special structure of the Maxwell system to show that the coefficients are uniquely determined.

The main contribution of this work is to show that the Carleman estimate approach to scalar partial data inverse problems introduced in [BU02, KSU07] can be carried over to the Maxwell system. A Carleman estimate approach for a different first order system, related to the Pauli Dirac operator, was presented in [ST09] where the method involved decoupling the Pauli Dirac operator into a second order differential operator with the Laplacian as its principal part. A suitable Carleman estimate was then applied to the decoupled equation to recover the coefficients. The partial data problem for Maxwell equations was also studied in [COS09] in the case when M is Euclidean and the inaccessible part of the boundary is a portion of a hyperplane or a sphere, following the scalar approach of [Is07]. Due to the partial symmetry of the domain one can reflect across the flat part of the boundary and reduce the problem to a full data problem. Partial data results are also known for a two-dimensional Maxwell system [IY15a] and for Maxwell equations in a waveguide [IY15b], based on the two-dimensional partial data results of [IUY10], and in the case where the parameters are known near the boundary [BMR16]. Recently, extensions of the Carleman estimate and reflection approaches for partial data problems were introduced in [KS13] and [IY13]; some of these extensions for the Maxwell system were considered in [IY14]. We also remark that the methods in the recent paper [DKLS16] might allow to relax to some extent the admissibility assumption in Theorem 1.1. Finally, we mention that while this article was being reviewed, the preprint [As16] containing an improvement of our main results appeared.

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2. REDUCTION TO THE HODGE LAPLACIAN

In this section we employ ideas from [OS96, OPS03, KSU09] to show that the problem reduces to constructing suitable solutions of the Hodge Dirac and Schrödinger operator. The notation follows closely [KSU09] and we refer to that article for more details. If E and H are complex 1-forms in M , we consider the graded form $X = \Phi + E + *H + *\Psi$ where Φ and Ψ are complex scalar functions. We write graded forms in the vector notation

$$X = \left(\begin{array}{c|c} \Phi & *H \\ \hline *\Psi & E \end{array} \right)^t.$$

The line denotes that the forms of even order and odd order are grouped together, which will result in a block structure for the equation. We define the following Hodge Dirac

operator acting on graded forms in M ,

$$P = \frac{1}{i}(d - \delta) = \frac{1}{i} \left(\begin{array}{c|c} & -\delta \\ \hline d & d \\ \hline d & -\delta \end{array} \right).$$

The Dirac equation which is most closely related to Maxwell is

$$(P + V)X = 0$$

where V is the operator acting on graded forms as the matrix operator

$$V = \left(\begin{array}{c|c} -\omega\mu & *D\alpha \wedge * \\ \hline -\omega\mu & *D\alpha \wedge * \\ D\beta\wedge & -\omega\varepsilon \\ \hline D\beta\wedge & -\omega\varepsilon \end{array} \right).$$

and where $D = \frac{1}{i}d$, $\alpha = \log \varepsilon$, and $\beta = \log \mu$. One easily sees that

$$(2.1) \quad \begin{cases} (P + V)X = 0, \\ X = (0 \quad *H \mid *0 \quad E)^t \end{cases} \Leftrightarrow \begin{cases} *dE = i\omega\mu H, \\ *dH = -i\omega\varepsilon E. \end{cases}$$

The advantage here is that P is an elliptic operator.

For the reduction to the Schrödinger equation with Hodge Laplacian, we consider rescaled solutions

$$Y = \left(\begin{array}{c|c} \mu^{1/2} & \\ \hline & \varepsilon^{1/2} \end{array} \right) X,$$

and henceforth relate X and Y in this way. We will write

$$Y = (Y^0 \quad Y^2 \mid Y^3 \quad Y^1)^t$$

where Y^k is the k -form part of the graded form Y . It is easy to see that

$$(P + V)X = 0 \iff (P + W)Y = 0$$

where W is the potential

$$W = -\kappa + \frac{1}{2} \left(\begin{array}{c|c} & *D\alpha \wedge * \\ \hline & *D\alpha \wedge * \quad -D\alpha\wedge \\ D\beta\wedge & \\ \hline D\beta\wedge & *D\beta \wedge * \end{array} \right),$$

and $\kappa = \omega(\varepsilon\mu)^{1/2}$.

The reason for using the rescaled solutions Y is that there is a good reduction of the Dirac equation $(P + W)Y = 0$ to a Schrödinger equation with no first order terms (see [KSU09, Lemma 3.1]).

Lemma 2.1. *Denote by $-\Delta = d\delta + \delta d$ the Hodge Laplacian acting on graded forms. We have*

$$\begin{aligned} (P + W)(P - W^t) &= -\Delta + Q, \\ (P - W^t)(P + W) &= -\Delta + Q', \\ (P + W^*)(P - \bar{W}) &= -\Delta + \hat{Q}, \end{aligned}$$

where Q , Q' , and \hat{Q} are continuous potentials (endomorphisms of ΛM) with

$$Q = -\kappa^2 + \frac{1}{2} \left(\begin{array}{cc|cc} \Delta\alpha + \frac{1}{2}\langle d\alpha, d\alpha \rangle & 0 & \bullet & \bullet \\ 0 & \bullet & \bullet & \bullet \\ \hline \bullet & \bullet & \Delta\beta + \frac{1}{2}\langle d\beta, d\beta \rangle & 0 \\ \bullet & \bullet & 0 & \bullet \end{array} \right),$$

$$Q' = -\kappa^2 - \frac{1}{2} \left(\begin{array}{cc|cc} \Delta\beta - \frac{1}{2}\langle d\beta, d\beta \rangle & 0 & 0 & 0 \\ \bullet & \bullet & \bullet & \bullet \\ \hline 0 & 0 & \Delta\alpha - \frac{1}{2}\langle d\alpha, d\alpha \rangle & 0 \\ \bullet & \bullet & \bullet & \bullet \end{array} \right),$$

and \bullet denote bounded coefficients involving derivatives of ε and μ up to second order.

Remarks. The coefficient functions denoted simply by a “bullet” above are not needed in our argument. They are easy to compute though if one so wishes. Also observe that due to this lemma, if Z is a solution of $(-\Delta + \hat{Q})Z = 0$, then $Y := (P - \bar{W})Z$ would solve $(P + W^*)Y = 0$. Finally, the precise form of \hat{Q} is not needed in this work, so we will not bother the reader by writing it out explicitly.

The above arguments can be used to reduce the Maxwell equations to the Hodge Dirac and Schrödinger equations. We now want to give a similar reduction on the level of boundary measurements. In fact, we will prove an integral identity showing - see Lemma 2.3 - that if the partial Cauchy data for two sets of coefficients coincide, then the difference of the corresponding potentials Q_j is orthogonal to a product of solutions to the Schrödinger and Dirac equations. This result is a version of [KSU09, Lemma 3.2] adapted to the partial data problem.

If U and V are graded forms on M , we use the notations

$$(U|V) = \int_M \sum_{j=0}^3 \langle U^j, \bar{V}^j \rangle dV, \quad (U|V)_{\partial M} = \int_{\partial M} \sum_{j=0}^3 \langle U^j, \bar{V}^j \rangle dS.$$

for the L^2 -inner products. We also use the same notation in the case when U and V are in appropriate dual Sobolev spaces. Note that in the second expression, U and V do not have to be graded forms of the manifold ∂M (for instance, the 1-form parts may have a normal component). We have the integration by parts formulas ([Ta99, Section 2.10])

$$(2.2) \quad (dU|V) = (\nu \wedge U|V)_{\partial M} + (U|\delta V),$$

$$(2.3) \quad (\delta U|V) = -(i_\nu U|V)_{\partial M} + (U|dV)$$

where ν is the outward normal and i_ν is the contraction with ν . Also, we define

$$H_\Delta = \{u \in L^2(M, \Lambda M); \Delta u \in L^2(M, \Lambda M)\}$$

and equip this with the natural graph norm

$$\|u\|_{H_\Delta}^2 = \|u\|^2 + \|\Delta u\|^2.$$

Here and in the rest of this paper, notation of the form $\|\cdot\|$ and $\|\cdot\|_M$ should be taken to indicate the relevant L^2 norm unless otherwise specified.

One feature of the method in this paper is that we need to deal with pairs of solutions, where one solution can be very nonsmooth (only H^{-1} in general). For this reason we will need the following result concerning smooth approximation and traces in the H_Δ space. The proof is given in the appendix.

Lemma 2.2. *The set $C^\infty(M, \Lambda M)$ is dense in H_Δ . The trace maps $u \mapsto (tu, t\delta u)$, $u \mapsto (t * u, t\delta * u)$, $u \mapsto (tu, tdu)$ and $u \mapsto (t * u, td * u)$, initially defined on $C^\infty(M, \Lambda M)$, have unique bounded extensions as maps*

$$H_\Delta \rightarrow H^{-1/2}(\partial M, \Lambda \partial M) \times H^{-3/2}(\partial M, \Lambda \partial M).$$

Given a pair of electromagnetic parameters $\{\varepsilon_j, \mu_j\}$, $j = 1, 2$, we define the corresponding matrix potentials V_j , Q_j and W_j by replacing ε and μ respectively with ε_j and μ_j in definitions above. The next result is the main integral identity.

Lemma 2.3. *Let $\varepsilon_j, \mu_j \in C^3(M)$, $j = 1, 2$, be complex functions with positive real parts, and assume that*

$$(2.4) \quad \varepsilon_1 = \varepsilon_2, \quad \mu_1 = \mu_2 \text{ to second order on } \partial M.$$

Let Γ_2 be any open subset of ∂M , and assume that $C_1^{\partial M, \Gamma_2} = C_2^{\partial M, \Gamma_2}$. Then

$$((Q_1 - Q_2)Z_1 | Y_2) = 0$$

for any graded forms $Z_1 \in H^3$ and $Y_2 \in H^{-1}$ such that

$$\begin{aligned} (-\Delta + Q_1)Z_1 &= 0, \\ Y_1 &= (P - W_1^t)Z_1 \text{ satisfies } Y_1^0 = Y_1^3 = 0, \end{aligned}$$

and

$$\begin{aligned} Y_2 &= (P - \bar{W}_2)Z_2 \text{ for some } Z_2 \in H_\Delta \text{ with } (-\Delta + \hat{Q}_2)Z_2 = 0, \\ &\text{supp}(tY_2) \subset \Gamma_2. \end{aligned}$$

Remark. By the assumptions for ε_j and μ_j , we have that $Q_1 - Q_2$ is C^1 and vanishes on ∂M . Thus $(Q_1 - Q_2)Z_1$ is H^1 and vanishes on ∂M , and the pairing $((Q_1 - Q_2)Z_1 | Y_2)$ where $Y_2 \in H^{-1}$ makes sense. Note also that since $Z_2 \in H_\Delta$, Lemma 2.2 guarantees that the boundary value tY_2 exists as an element of $H^{-3/2}$. Here, Z_1 and Y_1 need to be related to a solution for Maxwell, but Y_2 only needs to solve a Dirac equation. This simplifies the recovery of coefficients.

Proof of Lemma 2.3. By the discussion earlier in this section, if Y_1 is as stated, then $Y_1 \in H^2$ solves $(P + W_1)Y_1 = 0$ and gives rise to a H^2 solution $X_1 = \begin{pmatrix} 0 & *H_1 & | & 0 & E_1 \end{pmatrix}^t$ of the Maxwell equations $(P + V_1)X_1 = 0$. By the assumption on Cauchy data sets, there is $\tilde{X}_2 = \begin{pmatrix} 0 & *\tilde{H}_2 & | & 0 & \tilde{E}_2 \end{pmatrix}^t \in H^2$ satisfying

$$\begin{aligned} (P + V_2)\tilde{X}_2 &= 0, \\ t\tilde{E}_2|_{\partial M} &= tE_1|_{\partial M}, \quad t\tilde{H}_2|_{\Gamma_2} = tH_1|_{\Gamma_2}. \end{aligned}$$

Let $\tilde{Y}_2 \in H^2$ be the solution of $(P + W_2)\tilde{Y}_2 = 0$ corresponding to \tilde{X}_2 .

We first claim that the assumption on Cauchy data sets implies

$$(2.5) \quad ((W_1 - W_2)Y_1 | Y_2)_M = 0.$$

Here, Y_2 is only in H^{-1} since $Y_2 = (P - \bar{W}_2)Z_2$ where $Z_2 \in H_\Delta$, but the pairing is well defined because the quantity $(W_1 - W_2)Y_1$ is in H^1 with vanishing boundary values since $\varepsilon_1 = \varepsilon_2$ and $\mu_1 = \mu_2$ to second order on ∂M . To deal with the nonsmooth solution Y_2 , we use Lemma 2.2 and choose $Z_l \in C^\infty(M, \Lambda M)$ with $Z_l \rightarrow Z_2$ in H_Δ as $l \rightarrow \infty$, which implies in particular that $(P - \bar{W}_2)Z_l \rightarrow Y_2$ in H^{-1} . Noting that

$$(P + W_2)(Y_1 - \tilde{Y}_2) = (W_2 - W_1)Y_1,$$

we compute

$$\begin{aligned}
((W_2 - W_1)Y_1|Y_2)_M &= \lim_l ((W_2 - W_1)Y_1|(P - \bar{W}_2)Z_l)_M \\
&= \lim_l ((P + W_2)(Y_1 - \tilde{Y}_2)|(P - \bar{W}_2)Z_l)_M \\
&= \lim_l \left[\frac{1}{i} ((\nu \wedge + i_\nu)(Y_1 - \tilde{Y}_2)|(P - \bar{W}_2)Z_l)_{\partial M} + (Y_1 - \tilde{Y}_2|(P + W_2^*)(P - \bar{W}_2)Z_l)_M \right].
\end{aligned}$$

In the last line we used (2.2) and (2.3). For the last term, we have

$$(2.6) \quad (P + W_2^*)(P - \bar{W}_2)Z_l = (-\Delta + \hat{Q}_2)Z_l \rightarrow (-\Delta + \hat{Q}_2)Z_2 = 0 \text{ in } L^2$$

so the second term in brackets goes to zero as $l \rightarrow \infty$. For the boundary term, we claim that

$$(2.7) \quad t(Y_1 - \tilde{Y}_2) = 0 \text{ on } \partial M$$

$$(2.8) \quad t * (Y_1 - \tilde{Y}_2) = 0 \text{ on } \Gamma_2$$

$$(2.9) \quad tY_2 = 0 \text{ on } \Gamma_2^c.$$

Assuming these, one has $\nu \wedge (Y_1 - \tilde{Y}_2)|_{\partial M} = 0$ and $i_\nu(Y_1 - \tilde{Y}_2)|_{\Gamma_2} = 0$, and the boundary term becomes

$$\lim_l \frac{1}{i} (Y_1 - \tilde{Y}_2 | \nu \wedge (P - \bar{W}_2)Z_l)_{\Gamma_2^c}.$$

Here $t(P - \bar{W}_2)Z_l|_{\Gamma_2^c} \rightarrow tY_2|_{\Gamma_2^c} = 0$ in $H^{-3/2}$ by Lemma 2.2. Since $Y_1 - \tilde{Y}_2|_{\partial M}$ is in $H^{3/2}$, the limit of the pairing on Γ_2^c is well defined and in fact zero. This proves (2.5) modulo (2.7)–(2.9). To check (2.7), we see that

$$\nu \wedge (Y_1 - \tilde{Y}_2) = (0 \quad \varepsilon_1^{1/2} \nu \wedge (E_1 - \tilde{E}_2) \mid * \mu_1^{1/2} \langle \nu, H_1 - \tilde{H}_2 \rangle \quad 0).$$

From the hypotheses, $\nu \wedge (E_1 - \tilde{E}_2)|_{\partial M} = 0$, and using the surface divergence as in the proof of [KSU09, Lemma 3.2], we also have $\langle \nu, H_1 - \tilde{H}_2 \rangle|_{\partial M} = 0$. Therefore (2.7) is satisfied. Meanwhile,

$$\nu \wedge *(Y_1 - \tilde{Y}_2) = (0 \quad \mu_1^{1/2} \nu \wedge (H_1 - \tilde{H}_2) \mid * \varepsilon_1^{1/2} \langle \nu, E_1 - \tilde{E}_2 \rangle \quad 0).$$

From the hypotheses, $\nu \wedge (H_1 - \tilde{H}_2)|_{\Gamma_2} = 0$, and using the surface divergence as in the proof of [KSU09, Lemma 3.2], we also have $\langle \nu, E_1 - \tilde{E}_2 \rangle|_{\Gamma_2} = 0$. Therefore (2.8) is satisfied. The equation (2.9) follows straight from the definition of Y_2 .

Having proved (2.5), we next claim that

$$(2.10) \quad ((W_1 - W_2)Y_1|Y_2)_M = ((Q_1 - Q_2)Z_1|Y_2)_M.$$

Note that the pairing on the right is well defined, since $(Q_1 - Q_2)Z_1$ is in H^1 with vanishing boundary values. Since $Q_j = W_j \circ P - P \circ W_j^t - W_j W_j^t$, we have

$$\begin{aligned}
(W_1 - W_2)Y_1 &= (W_1 - W_2)(P - W_1^t)Z_1 \\
&= (Q_1 + P \circ W_1^t)Z_1 - (Q_2 + P \circ W_2^t + W_2(W_2^t - W_1^t))Z_1 \\
&= (Q_1 - Q_2)Z_1 + (P + W_2)W_1^t Z_1 - (P + W_2)W_2^t Z_1.
\end{aligned}$$

Choosing again $Z_l \in C^\infty(M, \Lambda M)$ with $Z_l \rightarrow Z_2$ in H_Δ , we have

$$\begin{aligned} ((W_1 - W_2)Y_1|Y_2)_M &= \lim_l ((W_1 - W_2)Y_1|(P - \bar{W}_2)Z_l)_M \\ &= \lim_l [((Q_1 - Q_2)Z_1|(P - \bar{W}_2)Z_l)_M + ((P + W_2)(W_1^t - W_2^t)Z_1|(P - \bar{W}_2)Z_l)_M]. \end{aligned}$$

The first term in brackets has limit $((Q_1 - Q_2)Z_1|Y_2)_M$. After integrating by parts using (2.2) and (2.3), the second term becomes

$$\lim_l \left[\frac{1}{i} ((\nu \wedge + i_\nu)(W_1^t - W_2^t)Z_1|(P - \bar{W}_2)Z_l)_{\partial M} + ((W_1^t - W_2^t)Z_1|(-\Delta + \hat{Q}_2)Z_l)_M \right].$$

Since $W_1^t = W_2^t$ at ∂M to first order, the boundary term vanishes. Also, by (2.6) the second term has limit zero. This proves (2.10), and combining this with (2.5) proves the lemma. \square

We now explain the connection between the Cauchy data set and the admittance map, which is a standard way for expressing boundary measurements for Maxwell equations. First we introduce some function spaces. The surface divergence of $f \in H^s(\partial M, \Lambda^1 \partial M)$ is given by

$$\text{Div}(f) = \langle d_{\partial M} f, dS \rangle$$

where dS is the volume form on ∂M . Define the spaces

$$\begin{aligned} H_{\text{Div}}^2(M) &= \{u \in H^2(M, \Lambda^1 M); \text{Div}(tu) \in H^{3/2}(\partial M)\}, \\ TH_{\text{Div}}^{3/2}(\partial M) &= \{f \in H^{3/2}(\partial M, \Lambda^1 \partial M); \text{Div}(f) \in H^{3/2}(\partial M)\}. \end{aligned}$$

By [KSU09, Theorem A.1 and subsequent remark], there is a discrete set $\Sigma \subset \mathbb{C}$ such that if $\omega \in \mathbb{C} \setminus \Sigma$, then for any $f \in TH_{\text{Div}}^{3/2}(\partial M)$ the Maxwell equations (1.1) have a unique solution $(E, H) \in H_{\text{Div}}^2(M) \times H_{\text{Div}}^2(M)$ satisfying $tE = f$ on ∂M . Assume that ω is not a resonant frequency, and let Γ_1 and Γ_2 be nonempty open subsets of ∂M . If $TH_{\text{Div},c}^{3/2}(\Gamma_1)$ consists of those elements in $TH_{\text{Div}}^{3/2}(\partial M)$ that are compactly supported in Γ_1 , we may define the partial admittance map

$$\Lambda^{\Gamma_1, \Gamma_2} : TH_{\text{Div},c}^{3/2}(\Gamma_1) \rightarrow TH_{\text{Div}}^{3/2}(\Gamma_2), \quad f \mapsto tH|_{\Gamma_2}$$

where (E, H) is the unique solution of (1.1) satisfying $tE = f$ on ∂M . Note also that any solution $(E, H) \in H^2 \times H^2$ of (1.1) is in $H_{\text{Div}}^2 \times H_{\text{Div}}^2$, by using the surface divergence as in [KSU09, Lemma 3.2].

Now if ω is not a resonant frequency, it is clear that C^{Γ_1, Γ_2} is the graph of $\Lambda^{\Gamma_1, \Gamma_2}$:

$$C^{\Gamma_1, \Gamma_2} = \{(f, \Lambda^{\Gamma_1, \Gamma_2} f); f \in TH_{\text{Div},c}^{3/2}(\Gamma_1)\}.$$

Thus C^{Γ_1, Γ_2} and $\Lambda^{\Gamma_1, \Gamma_2}$ are equivalent information in this case.

3. CARLEMAN ESTIMATES AND CGO SOLUTIONS

Let (M, g) be an admissible 3-manifold as in Definition 1.1, so that $M \subset \mathbb{R} \times M_0$ and $g = c(e \oplus g_0)$ where (M_0, g_0) is simple. Let (x_1, x') be global coordinates in $\mathbb{R} \times M_0$, and let $\varphi(x_1, x') = x_1$ be the natural LCW on M .

It is convenient to begin with a reduction to the case where $c \equiv 1$. This can be done using the following lemma from [KSU09].

Lemma 3.1. ([KSU09], Lemma 7.1) *Let (M, g) be a compact Riemannian 3-manifold with smooth boundary, and let ε and μ be C^2 functions on M with positive real part. Let c be any positive smooth function on M , and let $C_{g, \varepsilon, \mu}^{\Gamma_1, \Gamma_2}$ represent the Cauchy data for ε, μ with respect to the metric g . Then*

$$C_{cg, \varepsilon, \mu}^{\Gamma_1, \Gamma_2} = C_{g, c^{1/2}\varepsilon, c^{1/2}\mu}^{\Gamma_1, \Gamma_2}.$$

Proof. Lemma 7.1 in [KSU09] is actually phrased in terms of the admittance map, which requires the additional condition that solutions to the Maxwell equations are unique. However, the same proof works when the lemma is phrased in terms of the Cauchy data. The key fact is that for a k -form u ,

$$*_{cg}u = c^{3/2-k} *_g u,$$

so (E, H) satisfies (1.1) with metric cg and coefficients ε and μ if and only if (E, H) satisfies (1.1) with metric g and coefficients $c^{1/2}\varepsilon$ and $c^{1/2}\mu$. \square

Thus it suffices to prove Theorem 1.1 in the case where $c \equiv 1$. We will from now on assume that $g = e \oplus g_0$ where (M_0, g_0) is a simple 2-manifold, and that we are working with the LCW $\varphi(x) = x_1$. We will also assume that (\hat{M}_0, g_0) is another simple 2-manifold which is slightly larger than M_0 .

As indicated in Lemma 2.3, we will need to construct two types of complex geometrical optics solutions. First we need sufficiently regular solutions to the Maxwell equations, with no restrictions on their boundary values. By Section 2 it suffices to construct solutions Z to the equation $(-\Delta + Q)Z = 0$ in M such that $Y = (P - W^t)Z$ satisfies $Y^0 = Y^3 = 0$. Such solutions were obtained in [KSU09, Theorem 6.1a]. Second, we need solutions of the Dirac equation $(P + W^*)Y = 0$ satisfying $tY|_{\Gamma} = 0$. Here $\Gamma \subset \partial M$ is meant to be an open neighbourhood of $\partial M \setminus \tilde{F}$, where \tilde{F} is as in the statement of Theorem 1.1. By proper choice of Γ , we can ensure that

$$(3.1) \quad \langle d\varphi, \nu \rangle < 0 \text{ on } \bar{\Gamma}.$$

The next theorem states the existence and basic properties of both types of complex geometrical optics solutions.

Theorem 3.1. *Let $\varepsilon, \mu \in C^3(M)$ have positive real parts, and let W and Q be as in Section 2. Fix $p \in \hat{M}_0 \setminus M_0$, and let (r, θ) be polar normal coordinates in (\hat{M}_0, g_0) with center p , so that (x_1, r, θ) are global coordinates near M . Let s_0, t_0 be real constants, let $\lambda > 0$, and let $b(\theta) \in C^\infty(S^1)$.*

- (a) For $|\tau|$ sufficiently large and outside a countable subset of \mathbb{R} , one can construct solutions to the equation $(-\Delta + Q)Z = 0$ in M satisfying

$$(P + W)Y = 0, \quad Y = (P - W^t)Z, \\ Y^0 = Y^3 = 0$$

where $Z \in H^3(M, \Lambda M)$ has the form

$$(3.2) \quad Z = e^{-\tau(x_1 + ir)} |g(r, \theta)|^{-1/4} e^{i\lambda(x_1 + ir)} b(\theta) \begin{pmatrix} s_0 \\ 0 \\ t_0 * 1 \\ 0 \end{pmatrix} + e^{-\tau(x_1 + ir)} R$$

with $\|R\|_{H^s} \leq C\tau^{s-1}$ for $0 \leq s \leq 2$.

- (b) Let $\Gamma \subset \partial M$ be an open set so that (3.1) holds for $\varphi(x) = x_1$. For $\tau > 0$ sufficiently large, there exists a solution $Z \in H_\Delta$ to $(-\Delta + \hat{Q})Z = 0$ in M so that

$$(P + W^*)Y = 0, \quad Y = (P - \bar{W})Z,$$

and Y has the form

$$(3.3) \quad Y = e^{\tau(x_1+ir)} |g(r, \theta)|^{-1/4} \begin{pmatrix} s_0 \\ -is_0 dx^1 \wedge dr \\ t_0 * 1 \\ it_0 * dx^1 \wedge dr \end{pmatrix} + e^{\tau(x_1+ir)} (\tilde{R} + R')$$

with $\|\tilde{R}\|_{H^{-1}} \leq C\tau^{-3/2}$, $\|R'\|_{L^2} \leq C\tau^{-1/2}$, and Y satisfies the boundary condition

$$tY|_\Gamma = 0.$$

Proof of Theorem 3.1(a). It is proved in [KSU09, Theorem 6.1a] that there is a solution $Z \in H^2(M, \Lambda M)$ of this form. That theorem only states L^2 bounds for R , but the proof of the theorem relies on [KSU09, Proposition 5.1] which actually gives the H^s bounds for $0 \leq s \leq 2$. It remains to show that under the assumption $\varepsilon, \mu \in C^3(M)$ (which implies $Q \in C^1$) one has in fact $Z \in H^3(M, \Lambda M)$. This follows from interior elliptic regularity since the solution in [KSU09, Theorem 6.1a] is constructed in a neighbourhood of M , and Z satisfies $-\Delta Z = -QZ$ where $QZ \in H^1$ in this neighbourhood. \square

In the remainder of this section we will prove Theorem 3.1(b). This will be done by constructing a solution for the equation $(-\Delta + \hat{Q})Z = 0$ of the form

$$(3.4) \quad Z = e^{\tau x_1} (A + R), \quad tZ|_\Gamma = t(\delta + i\bar{W})Z|_\Gamma = 0,$$

for some appropriate amplitude A chosen so that

$$e^{-\tau x_1} (-\Delta + \hat{Q}) e^{\tau x_1} A = O_{L^2(M)}(1).$$

Then $Y = (P - \bar{W})Z$ solves $(P + W^*)Y = 0$ and satisfies the desired boundary conditions.

The remainder term R in (3.4) will be constructed using the following Carleman estimate for graded forms with suitable boundary conditions. If u is a graded form we write $u_\perp = \nu \wedge i_\nu u$ and $u_\parallel = u - u_\perp$. The proof of the following estimate was first given in [CST17]. Here $\|\cdot\| = \|\cdot\|_M$ and $\|\cdot\|_{\partial M}$ are the relevant L^2 norms in M and ∂M .

Proposition 3.2. ([CST17, Theorem 7.2]) *Let (M, g) be an admissible manifold as in Definition 1.1 and $\varphi(x) = x_1$ a LCW. Let \hat{Q} and σ be L^∞ endomorphisms on the space of graded forms. There is $C > 0$ such that for all smooth graded forms u satisfying*

$$te^{\tau\varphi} \delta e^{-\tau\varphi} u + t\sigma u_\perp|_\Gamma = tu|_\Gamma = 0,$$

$$u|_{\partial M \setminus \Gamma} = \nabla_\nu u|_{\partial M \setminus \Gamma} = 0,$$

and for all $\tau \geq C$, we have the following estimate

$$\tau \|u\| + \|\nabla u\| + \tau^{3/2} \|ti_\nu u\|_{\partial M} + \sqrt{\tau} \|\nabla' ti_\nu u\|_{\partial M} + \sqrt{\tau} \|t\nabla_\nu u_\parallel\|_{\partial M} \leq C \|e^{\tau\varphi} (-\Delta + \hat{Q}) e^{-\tau\varphi} u\|.$$

Proposition 3.2 relies on the fact that (3.1) holds on Γ – this is one of the key reasons for choosing \tilde{F} as we did in Theorem 1.1.

3.1. Construction of the amplitude. If \hat{Q} is an L^∞ endomorphism of ΛM , we look for a solution Z of

$$(-\Delta + \hat{Q})Z = 0 \quad \text{in } M$$

having the form

$$Z = e^{\tau x_1}(A + R)$$

where $\tau > 0$ is large. To this end we will construct an amplitude A such that

$$\|e^{-\tau x_1} \Delta e^{\tau x_1} A\|_M, \|A\|_M = O_{L^2(M)}(1).$$

Recall that since (M, g) is an admissible manifold and we have reduced to the case $c = 1$, the metric can be written globally as $g = e \oplus g_0$ for some conformal factor c and metric g_0 on the simple 2-dimensional manifold M_0 .

We now construct the amplitude A mentioned above.

Proposition 3.3. *For any smooth functions $b^j, b_r^j, b_\theta^j \in C^\infty(S^1)$ there exist graded forms $A = A^0 + A^1 + *B^1 + *B^0$ satisfying*

$$\|e^{-\tau x_1} \Delta e^{\tau x_1} A\|_M = O(1), \quad \|A\|_M = O(1)$$

of the form

$$\begin{aligned} A^0 &= |g_0|^{-1/4} e^{i\tau r} b^0(\theta), \\ A^1 &= |g_0|^{-1/4} e^{i\tau r} b^1(\theta) dx_1 + e^{i\tau r} (|g_0|^{-1/4} b_r^1(\theta) dr + |g_0|^{1/4} b_\theta^1(\theta) d\theta), \\ B^0 &= |g_0|^{-1/4} e^{i\tau r} b^3(\theta), \\ B^1 &= |g_0|^{-1/4} e^{i\tau r} b^2(\theta) dx_1 + e^{i\tau r} (|g_0|^{-1/4} b_r^2(\theta) dr + |g_0|^{1/4} b_\theta^2(\theta) d\theta). \end{aligned}$$

Proof. Any 1-form A^1 on M may be written as

$$\tilde{A}^1 = A_1^1 dx^1 + (A^1)'$$

with $A_1^1 = \langle A, dx^1 \rangle$ and $(A^1)'$ is a 1-form on M_0 that contains x_1 as a parameter. If A^0 is a 0-form and A^1 is a 1-form, we have

$$\begin{aligned} \Delta_g A^0 &= (\partial_1^2 + \Delta_{x'}) A^0, \\ \Delta_g A^1 &= (\Delta_g A_1^1) dx^1 + \partial_1^2 (A^1)' + \Delta_{x'} (A^1)' \end{aligned}$$

where $\Delta_{x'}$ is the Hodge Laplacian on (M_0, g_0) .

We compute the conjugated Hodge Laplacian acting on 0-forms and 1-forms,

$$\begin{aligned} e^{-\tau x_1} (-\Delta_g) e^{\tau x_1} A^0 &= (-\partial_1^2 - 2\tau \partial_1 - \tau^2 - \Delta_{x'}) A^0, \\ e^{-\tau x_1} (-\Delta_g) e^{\tau x_1} A^1 &= [(-\partial_1^2 - 2\tau \partial_1 - \tau^2 - \Delta_{x'}) A_1^1] dx^1 \\ &\quad + (-\nabla_{\partial_1}^2 - 2\tau \nabla_{\partial_1} - \tau^2 - \Delta_{x'}) (A^1)'. \end{aligned}$$

We now make the simplifying assumption that A^0 and A^1 only depend on x' . Then the above expressions simplify to

$$\begin{aligned} e^{-\tau x_1} (-\Delta_g) e^{\tau x_1} A^0 &= (-\Delta_{x'} - \tau^2) A^0, \\ e^{-\tau x_1} (-\Delta_g) e^{\tau x_1} A^1 &= [(-\Delta_{x'} - \tau^2) A_1^1] dx^1 + (-\Delta_{x'} - \tau^2) (A^1)'. \end{aligned}$$

It is enough to arrange that

$$(3.5) \quad \begin{aligned} \|(-\Delta_{x'} - \tau^2)A^0\|_{M_0} &= O(1), & \|A^0\|_{M_0} &= O(1), \\ \|(-\Delta_{x'} - \tau^2)A_1^1\|_{M_0} &= O(1), & \|A_1^1\|_{M_0} &= O(1), \\ \|(-\Delta_{x'} - \tau^2)(A^1)'\|_{M_0} &= O(1), & \|(A^1)'\|_{M_0} &= O(1). \end{aligned}$$

The construction for the scalar parts A^0 and A_1^1 is identical: we look for a WKB type quasimode

$$A^0 = e^{i\tau\psi} a$$

and note that

$$(-\Delta_{x'} - \tau^2)A^0 = e^{i\tau\psi} [\tau^2(|d\psi|^2 - 1)a - i\tau(2\nabla_{\nabla\psi}a + (\Delta\psi)a) - \Delta a].$$

In order to satisfy (3.5), we need ψ and a to satisfy the eikonal equation $|d\psi|^2 - 1 = 0$, and the transport equation $2\nabla_{\nabla\psi}a + (\Delta\psi)a = 0$. To find ψ , let (\hat{M}_0, g_0) be a slightly larger simple manifold than (M_0, g_0) , let $p \in \hat{M}_0 \setminus M_0$, let (r, θ) be polar normal coordinates in (\hat{M}_0, g_0) with center p , and choose

$$\psi(r, \theta) = r.$$

In these coordinates, the metric has the form

$$g_0(r, \theta) = \begin{pmatrix} 1 & 0 \\ 0 & m(r, \theta) \end{pmatrix}$$

for some smooth positive function m . It follows that

$$|d\psi|^2 = 1.$$

We also have $\nabla\psi = \partial_r$ and $\Delta\psi = \frac{1}{2}\partial_r(\log|g_0|)$ where $|g_0(r, \theta)|$ is the determinant of g_0 . The transport equation becomes

$$2\partial_r a + \frac{1}{2}\partial_r(\log|g_0|)a = 0$$

and it has the solution

$$a = |g_0|^{-1/4}b(\theta)$$

where b can be any function in $C^\infty(S^1)$. Therefore, setting $A_0 = |g_0|^{-1/4}e^{i\tau r}b^0(\theta)$, the first line of (3.5) is satisfied. A similar choice for A_1^1 , substituting b^1 for b^0 , would satisfy the second line of (3.5).

Moving on to $(A^1)'$, we look for an ansatz of the form

$$(A^1)' = e^{i\tau\psi} a', \quad a' = a_r dr + a_\theta d\theta.$$

We have

$$(-\Delta_{x'} - \tau^2)(A^1)' = e^{i\tau\psi} [\tau^2(|d\psi|^2 - 1)a' - i\tau[2\nabla_{\nabla\psi}a' + (\Delta\psi)a'] - \Delta a'.$$

Choosing the same function $\psi(r, \theta) = r$ as above we have $|d\psi|^2 = 1$. Since

$$\nabla_{\partial_r} dr = 0, \quad \nabla_{\partial_r} d\theta = -\frac{1}{2}\partial_r(\log|g_0|) d\theta,$$

the transport equations for a_r and a_θ become

$$\begin{aligned} 2\partial_r a_r + \frac{1}{2}\partial_r(\log|g_0|)a_r &= 0, \\ 2\partial_r a_\theta - \frac{1}{2}\partial_r(\log|g_0|)a_\theta &= 0. \end{aligned}$$

These have the solutions

$$a_r = |g_0|^{-1/4}b_r(\theta), \quad a_\theta = |g_0|^{1/4}b_\theta(\theta)$$

for any $b_\theta, b_r \in C^\infty(S^1)$. Setting $\tilde{A}' = e^{i\tau r}(|g_0|^{-1/4}b_r^1(\theta)dr + |g_0|^{1/4}b_\theta^1(\theta)d\theta)$, the third line of (3.5) is satisfied.

Now we can use the fact that $*$ commutes with Δ to construct B^0 and B^1 , with the same asymptotics, and finish the proof. \square

3.2. Construction of the remainder. We now construct the remainder term R with the appropriate boundary conditions and asymptotics in τ . More precisely we need to construct R to satisfy

$$e^{-\tau x_1}(-\Delta + \hat{Q})e^{\tau x_1}R = -e^{-\tau x_1}(-\Delta + \hat{Q})e^{\tau x_1}A$$

with boundary conditions

$$tR|_\Gamma = -tA|_\Gamma, \quad t(e^{-\tau x_1}\delta(e^{\tau x_1}R) + i\bar{W}R)|_\Gamma = -t(e^{-\tau x_1}\delta(e^{\tau x_1}A) + i\bar{W}A)|_\Gamma.$$

The first step in constructing such an R is the following solvability result similar to [CST17, Proposition 7.3], obtained from Proposition 3.2. Here we write $\Delta_\varphi = e^{\tau\varphi}\Delta e^{-\tau\varphi}$.

Proposition 3.4. *Let \hat{Q} and W be C^1 endomorphisms on M acting on graded forms. If $\tau > 0$ is large, for any graded forms $F \in L^2(M, \Lambda M)$ and $f_1, f_2 \in C^\infty(M, \Lambda M)$ there exists a solution $u \in L^2(M, \Lambda M) \cap e^{-\tau\varphi}H_\Delta$ to*

$$(3.6) \quad \begin{aligned} (-\Delta_{-\varphi} + \hat{Q})u &= F \\ tu|_\Gamma &= tf_1|_\Gamma, \\ (e^{-\tau\varphi}t\delta e^{\tau\varphi}u + tWu_\perp)|_\Gamma &= (e^{-\tau\varphi}t\delta e^{\tau\varphi}f_2 + tW(f_2)_\perp)|_\Gamma \end{aligned}$$

and which satisfies the estimate

$$(3.7) \quad \|u\|_M \leq \frac{C}{\tau} \left(\|F\|_M + \frac{1}{\sqrt{\tau}} \|e^{-\tau\varphi}t\delta e^{\tau\varphi}f_2 + tW(f_2)_\perp\|_{\partial M} + \sqrt{\tau} \|tf_1\|_{\partial M} \right).$$

Note that the boundary conditions in (3.6) make sense because $u \in e^{-\tau\varphi}H_\Delta$ guarantees that the traces in (3.6) are well defined; see Lemma 2.2.

Proof. Let H be the subspace of $L^2(M, \Lambda M)$ defined by

$$H := \{(-\Delta_\varphi + \hat{Q}^*)v \mid v \in C^\infty, tv = 0, e^{\tau\varphi}t\delta(e^{-\tau\varphi}v) = -ti_\nu W^*i_\nu v, v|_{\Gamma^c} = \nabla_N v|_{\Gamma^c} = 0\}$$

Define on H the linear operator

$$L((-\Delta_\varphi + \hat{Q}^*)v) := (v|_F)_M + (e^{\tau\varphi}i_\nu de^{-\tau\varphi}v|tf_1)_\Gamma + (i_\nu v|e^{-\tau\varphi}t\delta(e^{\tau\varphi}f_2) + tW(f_2)_\perp)_\Gamma.$$

We see that L is well defined and bounded, since by Proposition 3.2 and by the fact that $ti_\nu dv = -\nabla' ti_\nu v$ for v as above [CST17, Lemma 3.4], the quantity $|L(w)|$ is bounded by

$$\begin{aligned} & \|v\|_M \|F\|_M + \frac{\|te^{\tau\varphi} i_\nu d e^{-\tau\varphi} v\|_\Gamma}{\sqrt{\tau}} \sqrt{\tau} \|tf_1\|_\Gamma + \sqrt{\tau} \|i_\nu v\|_\Gamma \frac{\|e^{-\tau\varphi} t\delta(e^{\tau\varphi} f_2) + tW(f_2)_\perp\|_\Gamma}{\sqrt{\tau}} \\ & \leq \frac{C}{\tau} \|(-\Delta_\varphi + \hat{Q}^*)v\|_M \left(\|F\|_M + \sqrt{\tau} \|tf_1\|_\Gamma + \frac{1}{\sqrt{\tau}} \|e^{-\tau\varphi} t\delta(e^{\tau\varphi} f_2) + tW(f_2)_\perp\|_\Gamma \right). \end{aligned}$$

By Hahn-Banach, there exists $u \in L^2$ with the estimate

$$\|u\|_M \leq \frac{C}{\tau} \left(\|F\|_M + \sqrt{\tau} \|tf_1\|_\Gamma + \frac{1}{\sqrt{\tau}} \|e^{-\tau\varphi} t\delta(e^{\tau\varphi} f_2) + tW(f_2)_\perp\|_\Gamma \right)$$

satisfying

$$\begin{aligned} ((-\Delta_\varphi + \hat{Q}^*)v|u) &= L((-\Delta_\varphi + \hat{Q}^*)v) \\ &= (v|F)_M + (i_\nu d(e^{-\tau\varphi} v)|te^{\tau\varphi} f_1)_\Gamma + (i_\nu(e^{-\tau\varphi} v)|t\delta(e^{\tau\varphi} f_2) + e^{\tau\varphi} tW(f_2)_\perp)_\Gamma \end{aligned}$$

for all v as in the definition of H .

We consider first those v which are compactly supported in the interior of M to see that u solves $(-\Delta_{-\varphi} + \hat{Q})u = F$. Thus in particular $u \in e^{-\tau\varphi} H_\Delta$, so the traces of $e^{\tau\varphi} u$ are well defined by Lemma 2.2. Now we can integrate by parts using (2.2), (2.3), the relation $e^{\tau\varphi} t\delta(e^{-\tau\varphi} v) = -i_\nu W^* i_\nu v$, and the fact that v vanishes to first order on Γ^c , to see that on the boundary u satisfies

$$\begin{aligned} & (te^{\tau\varphi} u|i_\nu d(e^{-\tau\varphi} v))_\Gamma + (te^{-\tau\varphi} \delta(e^{\tau\varphi} u)|ti_\nu v)_\Gamma + (i_\nu u|i_\nu W^* i_\nu v)_\Gamma \\ &= (te^{\tau\varphi} f_1|i_\nu d(e^{-\tau\varphi} v))_\Gamma + (te^{-\tau\varphi} \delta(e^{\tau\varphi} f_2) + tW(f_2)_\perp|ti_\nu v)_\Gamma \end{aligned}$$

which becomes

$$\begin{aligned} & (te^{\tau\varphi} u|i_\nu d(e^{-\tau\varphi} v))_\Gamma + (te^{-\tau\varphi} \delta(e^{\tau\varphi} u) + tWu_\perp|ti_\nu v)_\Gamma \\ &= (te^{\tau\varphi} f_1|i_\nu d(e^{-\tau\varphi} v))_\Gamma + (te^{-\tau\varphi} \delta(e^{\tau\varphi} f_2) + tW(f_2)_\perp|ti_\nu v)_\Gamma \end{aligned}$$

for all v as in the definition of H . Now for all $\psi \in \Omega^{k+1}(M)$ with $\text{supp}(i_\nu \psi) \subset \Gamma$, we may choose $v \in \Omega^k(M)$ such that

$$v|_{\partial M} = 0, \quad t\delta(e^{-\tau\varphi} v) = 0, \quad i_\nu d(e^{-\tau\varphi} v) = i_\nu \psi$$

(See Lemma 3.3.2 of [Sc95].) A brief computation shows that v vanishes to first order on Γ^c and can be used in the above identity, which gives

$$(te^{\tau\varphi} u|i_\nu \psi)_\Gamma = (te^{\tau\varphi} f_1|i_\nu \psi)_\Gamma$$

whenever $\text{supp}(i_\nu \psi) \subset \Gamma$. Therefore we see that $te^{\tau\varphi} u = te^{\tau\varphi} f_1$ on Γ and

$$(te^{-\tau\varphi} \delta(e^{\tau\varphi} u) + tWu_\perp|ti_\nu v)_\Gamma = (te^{-\tau\varphi} \delta(e^{\tau\varphi} f_2) + tW(f_2)_\perp|ti_\nu v)_\Gamma$$

for all v as in the definition of H . Applying an argument similar to [Sc95, Lemma 3.3.2] (see also Lemma 5.3) again allows us to conclude that

$$e^{-\tau\varphi} t\delta(e^{\tau\varphi} u) + tWu_\perp = te^{-\tau\varphi} \delta(e^{\tau\varphi} f_2) + tW(f_2)_\perp$$

on Γ . □

This gives the following existence result for complex geometrical optics solutions with boundary conditions.

Proposition 3.5. *Let \hat{Q} and W be C^1 endomorphisms acting on graded forms. For $\tau > 0$ sufficiently large there exist solutions to the equation*

$$(-\Delta + \hat{Q})Z = 0$$

of the form

$$Z = e^{\tau x_1}(A + R),$$

with boundary conditions

$$tZ|_{\Gamma} = 0, \quad t(\delta Z + i\bar{W}Z_{\perp})|_{\Gamma} = 0$$

where A is as in Proposition 3.3 with $R \in L^2 \cap e^{-\tau x_1}H_{\Delta}$ satisfying the estimate $\|R\| \leq \frac{C}{\sqrt{\tau}}$.

Proof. By Proposition 3.3, the L^2 norm of $F := e^{-\tau x_1}(-\Delta + \hat{Q})e^{\tau x_1}A$ is $O(1)$ as $\tau \rightarrow \infty$. Now apply Proposition 3.4 to obtain $R \in L^2 \cap e^{-\tau x_1}H_{\Delta}$ solving

$$\begin{aligned} e^{-\tau x_1}(-\Delta + \hat{Q})e^{\tau x_1}R &= -F, \\ tR|_{\Gamma} &= -tA|_{\Gamma}, \\ t(e^{-\tau x_1}\delta e^{\tau x_1}R + i\bar{W}R_{\perp})|_{\Gamma} &= -t(e^{-\tau x_1}\delta e^{\tau x_1}A + i\bar{W}A_{\perp})|_{\Gamma} \end{aligned}$$

with the estimates

$$\begin{aligned} \|R\|_M &\leq \frac{C}{\tau} \left(\|F\|_M + \sqrt{\tau}\|tA\|_{\partial M} + \frac{1}{\sqrt{\tau}}\|e^{-\tau x_1}t\delta e^{\tau x_1}A + ti\bar{W}A_{\perp}\|_{\partial M} \right) \\ &\leq \frac{C}{\sqrt{\tau}}. \end{aligned}$$

□

Proof of Theorem 3.1(b). By Proposition 3.5, there is a solution $Z_0 \in H_{\Delta}$ of the equation $(-\Delta + \hat{Q})Z_0 = 0$ satisfying

$$Z_0 = e^{\tau x_1}(A + R), \quad tZ_0|_{\Gamma} = t(\delta + i\bar{W})Z_0|_{\Gamma} = 0$$

where the amplitude A is chosen to be

$$A = e^{i\tau r}|g|^{-1/4}b(\theta)\zeta_2$$

with $\zeta_2 = -i(s_0 dx^1 + t_0 * dx^1)$ where s_0, t_0 are real constants.

We define $Z := \frac{1}{\tau}Z_0$, so Z will solve the same equation as Z_0 with the same boundary conditions. Then $Y := (P - \bar{W})Z$ will solve $(P + W^*)Y = 0$ in the sense of distributions. Since $tZ|_{\Gamma} = 0$, we have $Z|_{\Gamma} = Z_{\perp}|_{\Gamma}$ and $t\delta Z|_{\Gamma} = 0$ by Lemma 5.1. Consequently

$$tY|_{\Gamma} = \frac{1}{i}(t\delta Z - t\delta Z - it\bar{W}Z)|_{\Gamma} = -\frac{1}{i}t(\delta Z + i\bar{W}Z_{\perp})|_{\Gamma} = 0.$$

Finally, we observe that

$$\begin{aligned} Y &= \frac{1}{\tau}(P - \bar{W})(e^{\tau(x_1+ir)}|g|^{-1/4}b(\theta)\zeta + e^{\tau x_1}R) \\ &= \frac{1}{\tau}P(e^{\tau(x_1+ir)}|g|^{-1/4}b(\theta)\zeta) - \frac{1}{\tau}\bar{W}e^{\tau(x_1+ir)}|g|^{-1/4}b(\theta)\zeta + \frac{1}{\tau}[P, e^{\tau x_1}]R \\ &\quad + e^{\tau x_1} \underbrace{\frac{1}{\tau}(P - \bar{W})(R)}_{\hat{R}}. \end{aligned}$$

A direct computation (see [KSU09, Proof of Theorem 6.1(b)]) gives the result. \square

4. RECOVERY OF COEFFICIENTS

Here we use the solutions constructed in Theorem 3.1 to recover the coefficients of the Maxwell equations from partial boundary measurements. The argument is essentially the same as in [KSU09].

Proof of Theorem 1.1. Let Z_1 be the solution to $(-\Delta + Q_1)Z_1 = 0$ constructed in Theorem 3.1(a) having the form

$$Z_1 = e^{-\tau(x_1+ir)}|g|^{-1/4}e^{i\lambda(x_1+ir)}b(\theta) \begin{pmatrix} s_0 \\ 0 \\ t_0 * 1 \\ 0 \end{pmatrix} + e^{-\tau(x_1+ir)}R.$$

Choose an open set $F_1 \subset \partial M$ such that $F_\varphi \subset F_1 \subset \subset \tilde{F}$, and let $\Gamma = \partial M \setminus \bar{F}_1$. Then (3.1) holds, and we can take Y_2 to be the solution to $(P + W_2^*)Y_2 = 0$ constructed in Theorem 3.1(b), satisfying $tY_2|_\Gamma = 0$ and having the form

$$Y_2 = e^{\tau(x_1-ir)}|g|^{-1/4} \begin{pmatrix} s_0 \\ is_0 dx^1 \wedge dr \\ t_0 * 1 \\ -it_0 * dx^1 \wedge dr \end{pmatrix} + e^{\tau(x_1-ir)}(\tilde{R} + R').$$

(In fact we solve $(P - \bar{W}_2^*)Y = 0$ where Y is as in Theorem 3.1(b), and we let $Y_2 = \bar{Y}$.) Here s_0, t_0, λ are any real constants with $\lambda > 0$, $b(\theta)$ is any function in $C^\infty(S^1)$, and $\tau > 0$ can be taken arbitrarily large.

By taking $\Gamma_2 = \tilde{F}$ in Lemma 2.3 we have $\text{supp}(tY_2) \subset \Gamma_2$. Thus we have the integral identity

$$((Q_1 - Q_2)Z_1|Y_2) = \int_M \langle (Q_1 - Q_2)Z_1|\bar{Y}_2 \rangle dV = 0.$$

Writing this out term by term we have

$$(4.1) \quad \int_M \left\langle (Q_1 - Q_2)e^{i\lambda(x_1+ir)}b(\theta) \begin{pmatrix} s_0 \\ 0 \\ t_0 * 1 \\ 0 \end{pmatrix} \middle| \begin{pmatrix} s_0 \\ -is_0 dx^1 \wedge dr \\ t_0 * 1 \\ it_0 * dx^1 \wedge dr \end{pmatrix} \right\rangle \\ + \int_M \left\langle (Q_1 - Q_2)R|(\tilde{R} + R') \right\rangle |g|^{1/2} + \int_M \left\langle (Q_1 - Q_2)e^{i\lambda(x_1+ir)}b(\theta) \begin{pmatrix} s_0 \\ 0 \\ t_0 * 1 \\ 0 \end{pmatrix} \middle| \tilde{R} + R' \right\rangle |g|^{1/4} \\ + \int_M \left\langle (Q_1 - Q_2)R \middle| \begin{pmatrix} s_0 \\ -is_0 dx^1 \wedge dr \\ t_0 * 1 \\ it_0 * dx^1 \wedge dr \end{pmatrix} \right\rangle |g|^{1/4} = 0.$$

Here all integrals are with respect to the 3-form $|g|^{-1/2} dV = dx_1 dr d\theta$.

We will argue that all terms in (4.1) except for the first one vanish as $\tau \rightarrow \infty$. Indeed, by Theorem 3.1, $\|R\|_{H^s} \leq C\tau^{s-1}$, $\|\tilde{R}\|_{H^{-1}} \leq C\tau^{-3/2}$, and $\|R'\|_{L^2} \leq C\tau^{-1/2}$. Using these

estimates for $s = 1$ and $s = 0$ we see that the second term is $O(\tau^{-3/2})$. Here we use that $\varepsilon_1 = \varepsilon_2$ and $\mu_1 = \mu_2$ to second order on ∂M , to have $Q_1 = Q_2$ on ∂M . By the same estimates for $s = 0$ we see that the third term is $O(\tau^{-1/2})$ and the fourth term is $O(\tau^{-1})$. Thus taking $\tau \rightarrow \infty$ in (4.1) implies that

$$\int_M \left\langle (Q_1 - Q_2) \begin{pmatrix} s_0 \\ 0 \\ t_0 * 1 \\ 0 \end{pmatrix} \middle| \begin{pmatrix} s_0 \\ -is_0 dx^1 \wedge dr \\ t_0 * 1 \\ it_0 * dx^1 \wedge dr \end{pmatrix} \right\rangle e^{i\lambda(x_1 + ir)} b(\theta) dx_1 dr d\theta = 0$$

for all real s_0, t_0, λ with $\lambda > 0$, and all smooth functions $b(\theta) \in C^\infty(S^1)$.

The rest of the proof is identical with the proof of Theorem 1.1 in [KSU09]; we give the argument for completeness. Let q_α and q_β be the elements of $Q_1 - Q_2$, interpreted as a 8×8 matrix, which correspond to the (1, 1)th and (5, 5)th elements, respectively. By Lemma 2.1

$$\begin{aligned} q_\alpha &= \frac{1}{2}\Delta(\alpha_1 - \alpha_2) + \frac{1}{4}\langle d\alpha_1, d\alpha_1 \rangle - \frac{1}{4}\langle d\alpha_2, d\alpha_2 \rangle - \omega^2(\varepsilon_1\mu_1 - \varepsilon_2\mu_2), \\ q_\beta &= \frac{1}{2}\Delta(\beta_1 - \beta_2) + \frac{1}{4}\langle d\beta_1, d\beta_1 \rangle - \frac{1}{4}\langle d\beta_2, d\beta_2 \rangle - \omega^2(\varepsilon_1\mu_1 - \varepsilon_2\mu_2). \end{aligned}$$

With the two choices $(s_0, t_0) = (1, 0)$ and $(s_0, t_0) = (0, 1)$, the special form of Q_1 and Q_2 in Lemma 2.1 shows that we obtain the two identities

$$\begin{aligned} \int_M e^{i\lambda(x_1 + ir)} b(\theta) q_\alpha(x) dx_1 dr d\theta &= 0, \\ \int_M e^{i\lambda(x_1 + ir)} b(\theta) q_\beta(x) dx_1 dr d\theta &= 0. \end{aligned}$$

Let $T = \mathbb{R} \times M_0$. Since $Q_1|_{\partial M} = Q_2|_{\partial M}$, the zero extensions of q_α and q_β to $T \setminus M$ are continuous functions and the integrals above may be taken over T . Varying $b(\theta)$, it follows that for all θ we have

$$\int_0^\infty e^{-\lambda r} \left[\int_{-\infty}^\infty e^{i\lambda x_1} q_\alpha(x_1, r, \theta) dx_1 \right] dr = 0$$

and similarly for q_β . Since (r, θ) are polar normal coordinates in M_0 , the curves $r \mapsto (r, \theta)$ are geodesics in M_0 . Denoting the expression in brackets by $f_\alpha(r, \theta)$ and varying the center of polar normal coordinates in Theorem 3.1 and varying θ , we obtain that

$$\int_0^\infty f_\alpha(\gamma(r)) \exp \left[- \int_0^r \lambda ds \right] dr = 0$$

for all geodesics γ in M_0 which begin and end at points of ∂M_0 . This shows the vanishing of the geodesic ray transform of the function f_α with constant attenuation $-\lambda$. Since this transform is injective on simple two-dimensional manifolds [SU11], we have $f_\alpha \equiv 0$ for all λ . Thus

$$\int_{-\infty}^\infty e^{i\lambda x_1} q_\alpha(x_1, r, \theta) dx_1 = 0$$

for all $\lambda > 0, r$ and θ . Uniqueness for the Fourier transform in x_1 shows that $q_\alpha \equiv 0$ in M . We obtain $q_\beta \equiv 0$ in M by the exact same argument.

We have arrived at the following two equations in M :

$$\begin{aligned} -\frac{1}{2}\Delta(\alpha_1 - \alpha_2) - \frac{1}{4}\langle d(\alpha_1 + \alpha_2), d(\alpha_1 - \alpha_2) \rangle + \omega^2(\varepsilon_1\mu_1 - \varepsilon_2\mu_2) &= 0, \\ -\frac{1}{2}\Delta(\beta_1 - \beta_2) - \frac{1}{4}\langle d(\beta_1 + \beta_2), d(\beta_1 - \beta_2) \rangle + \omega^2(\varepsilon_1\mu_1 - \varepsilon_2\mu_2) &= 0. \end{aligned}$$

Let $u = (\varepsilon_1/\varepsilon_2)^{1/2}$ and $v = (\mu_1/\mu_2)^{1/2}$. Then $\frac{1}{2}(\alpha_1 - \alpha_2) = \log u$ and $\frac{1}{2}(\beta_1 - \beta_2) = \log v$, and the equations become

$$\begin{aligned} -\Delta(\log u) - (\varepsilon_1\varepsilon_2)^{-1/2}\langle d(\varepsilon_1\varepsilon_2)^{1/2}, d(\log u) \rangle + \omega^2(\varepsilon_1\mu_1 - \varepsilon_2\mu_2) &= 0, \\ -\Delta(\log v) - (\mu_1\mu_2)^{-1/2}\langle d(\mu_1\mu_2)^{1/2}, d(\log v) \rangle + \omega^2(\varepsilon_1\mu_1 - \varepsilon_2\mu_2) &= 0. \end{aligned}$$

Multiplying the first equation by $(\varepsilon_1\varepsilon_2)^{1/2}$ and the second by $(\mu_1\mu_2)^{1/2}$ and using that $\delta(adw) = -a\Delta w - \langle da, dw \rangle$, we obtain the equations

$$\begin{aligned} \delta((\varepsilon_1\varepsilon_2)^{1/2}d(\log u)) + \omega^2(\varepsilon_1\varepsilon_2)^{1/2}\varepsilon_2\mu_2(u^2v^2 - 1) &= 0, \\ \delta((\mu_1\mu_2)^{1/2}d(\log v)) + \omega^2(\mu_1\mu_2)^{1/2}\varepsilon_2\mu_2(u^2v^2 - 1) &= 0. \end{aligned}$$

Since $d(\log u) = u^{-1}du$, $d(\log v) = v^{-1}dv$, we see that u and v satisfy the semilinear elliptic system

$$\begin{aligned} \delta(\varepsilon_2du) + \omega^2\varepsilon_2^2\mu_2(u^2v^2 - 1)u &= 0, \\ \delta(\mu_2dv) + \omega^2\varepsilon_2\mu_2^2(u^2v^2 - 1)v &= 0. \end{aligned}$$

The assumptions on ε_j and μ_j on ∂M imply that $u = v = 1$ and $\partial_\nu u = \partial_\nu v = 0$ on ∂M . Also, the above equations imply that the pair $(\tilde{u}, \tilde{v}) = (1, 1)$ is a solution of the semilinear system in all of M . Unique continuation holds for this system (see for instance [KSU09, Appendix B]), and we obtain $u \equiv 1$ and $v \equiv 1$ in M . This proves that $\varepsilon_1 \equiv \varepsilon_2$ and $\mu_1 \equiv \mu_2$ in M as required. \square

Proof of Theorem 1.2. Assume for simplicity that $\bar{\Omega} \subset \{x_3 > 0\}$ and that $x_0 = 0$. Let $\varphi(x) = \log |x|$ be an LCW in \mathbb{R}^3 . As explained in [KS13, Section 3B], if one chooses new coordinates

$$y_1 = \log |x|, \quad y' = \frac{x}{|x|}$$

where $y' \in S_+^2 = \{x \in S^2; x_3 > 0\}$, then the LCW becomes $\varphi(y) = y_1$ and the Euclidean metric becomes

$$g(y) = c(y) \begin{pmatrix} 1 & 0 \\ 0 & g_0(y') \end{pmatrix}$$

where g_0 is the standard metric on the sphere S^2 . Since manifolds $\{x \in S_+^2; x_3 \geq c > 0\}$ are simple, we have reduced matters to Theorem 1.1. This proves Theorem 1.2. \square

5. APPENDIX

For a compact oriented smooth manifold M with smooth boundary, define

$$H_\Delta = \{u \in L^2(M, \Lambda M) \mid \Delta u \in L^2(M, \Lambda M)\},$$

where Δ is the Hodge Laplacian. We let

$$\|u\|_{H_\Delta}^2 = \|u\|_{L^2}^2 + \|\Delta u\|_{L^2}^2.$$

In this appendix, we will prove the following trace theorem for H_Δ .

Lemma 5.1. *The relative and absolute trace maps $u \mapsto (tu, t\delta u)$ and $u \mapsto (t * u, t\delta * u)$ initially defined on the space $C^\infty(M, \Lambda M)$ have extensions as bounded linear maps from H_Δ to $H^{-1/2}(\partial M, \Lambda\partial M) \times H^{-3/2}(\partial M, \Lambda\partial M)$. The set $C^\infty(M, \Lambda M)$ is dense in H_Δ and*

thus these extensions are unique. The trace maps $u \mapsto (tu, tdu)$ and $u \mapsto (t * u, td * u)$ have extensions to the same spaces, and one has for any $u \in H_\Delta$

$$(5.1) \quad t(du) = d_{\partial M}(tu),$$

$$(5.2) \quad t(d * u) = d_{\partial M}(t * u).$$

Lemma 2.2 follows immediately from this result. To begin the proof of Lemma 5.1, we will need the following two lemmas. Here ν is meant to signify the normal vector field on ∂M , extended into M in a standard way.

Lemma 5.2. *Let $j \geq 0$. For any $v \in H^j(M, \Lambda M)$ and $f, h \in H^{j+3/2}(\partial M, \Lambda \partial M)$, there is a unique $u \in H^{j+2}(M, \Lambda M)$ that solves*

$$\Delta u = v \text{ in } M, \quad tu = f, \quad t * u = h.$$

One has $\|u\|_{H^{j+2}} \leq C(\|v\|_{H^j} + \|f\|_{H^{j+3/2}} + \|h\|_{H^{j+3/2}})$.

Proof. This is [Sc95, Theorem 3.4.10]. \square

Lemma 5.3. *For any $(f, h) \in H^{\frac{3}{2}}(\partial M, \Lambda \partial M) \times H^{\frac{1}{2}}(\partial M, \Lambda \partial M)$, there is $v \in H^2(M, \Lambda M)$ such that*

$$\|v\|_{H^2} \lesssim \|f\|_{H^{3/2}} + \|h\|_{H^{1/2}}$$

and $ti_\nu v = f$, $ti_\nu * v = ti_\nu d * v = 0$, and $ti_\nu dv = h$.

Proof. Let first $f = 0$. Any boundary point has a neighborhood $U \subset M$ where boundary normal coordinates (x_1, \dots, x_n) exist, so that the boundary is given by $\{x_n = 0\}$ and x_n coincides with the direction of ν . We first define v in U using these coordinates. Then v is represented as a sum of objects of the form

$$v_{i_1, \dots, i_k} dx_{i_1} \wedge \dots \wedge dx_{i_k}$$

for $1 \leq i_k \leq n$, where v_{i_1, \dots, i_k} are functions. Moreover, h can be represented as a sum of objects of the form

$$h_{i_1, \dots, i_k} dx_{i_1} \wedge \dots \wedge dx_{i_k},$$

for $1 \leq i_k < n$.

We will define v_{i_1, \dots, i_k} to be identically zero if any $i_k = n$, and choose the remaining v_{i_1, \dots, i_k} to be functions such that

$$\begin{aligned} v_{i_1, \dots, i_k} &= 0 \text{ and} \\ \partial_\nu v_{i_1, \dots, i_k} &= h_{i_1, \dots, i_k} \end{aligned}$$

on $U \cap \partial M$. Observe that when chosen this way all components of the form $v_{i_1, \dots, i_k} dx_{i_1} \wedge \dots \wedge dx_{i_k}$ vanish on the boundary and $i_\nu v_{i_1, \dots, i_k} dx_{i_1} \wedge \dots \wedge dx_{i_k}$ vanish in a neighborhood of the boundary. We can arrange for $\|v_{i_1, \dots, i_k}\|_{H^2(U)} \lesssim \|h_{i_1, \dots, i_k}\|_{H^{1/2}(U \cap \partial M)}$, as an inequality for functions. Using a suitable partition of unity we can define v near ∂M and extend smoothly as an element $v \in H^2(M, \Lambda M)$, and so we will obtain the inequality

$$\|v\|_{H^2(M, \Lambda M)} \lesssim \|h\|_{H^{1/2}(\partial M, \Lambda \partial M)}$$

for v and h as forms. Note that our definition is made to guarantee that $v|_{\partial M} = 0$, $t\nabla_\nu v = h$, and $i_\nu v = 0$ in a neighborhood of the boundary of M . Now by Lemma 3.4 of [CST17],

$$ti_\nu dv = t\nabla_\nu v_\parallel + Stv_\parallel - d'ti_\nu v$$

where S is a bounded endomorphism on ΛM and d' denotes the exterior derivative on ∂M . The last two terms vanish because v is zero on ∂M .

Then we are left with

$$ti_\nu dv = t\nabla_\nu v_{\parallel} = t\nabla_\nu v = h.$$

Moreover, $ti_\nu v = ti_\nu * v = 0$ since v is zero on the boundary of M , and $ti_\nu d * v = 0$ since

$$ti_\nu d * v = t\nabla_\nu (*v)_{\parallel} + St(*v)_{\parallel} - d'ti_\nu * v.$$

and $(*v)_{\parallel} = *v_{\perp} = 0$ in a neighbourhood of the boundary. This proves the lemma when $f = 0$. The next step is to consider the case where f is nonzero but $h = 0$. An argument in boundary normal coordinates similar to the one above yields a form v with $ti_\nu v = f$, $ti_\nu * v = ti_\nu d * v = ti_\nu dv = 0$, and $\|v\|_{H^2} \lesssim \|f\|_{H^{3/2}}$. Combining these two cases proves the lemma. \square

Proof of Lemma 5.1. Let $u, v \in H^2(M, \Lambda M)$. Then using the integration by parts formulas (2.2) and (2.3) we get the identity

$$(5.3) \quad (u|\Delta v)_M - (\Delta u|v)_M = (u|i_\nu dv - \nu \wedge \delta v)_{\partial M} + (\nu \wedge \delta u - i_\nu du|v)_{\partial M}.$$

We can rewrite the boundary terms in terms of the relative and absolute boundary values of u , by using the following formulas valid for any 1-form ξ and any k -form η (see for example [CST17]):

$$i_\xi \eta = (-1)^{n(k-1)} * \xi \wedge * \eta, \quad ** \eta = (-1)^{k(n-k)} \eta, \quad \delta \eta = (-1)^{(k-1)(n-k)+1} * d * \eta, \\ \langle *u, *v \rangle = \langle u, v \rangle, \quad \langle \xi \wedge u, v \rangle = \langle u, i_\xi v \rangle, \quad \langle i_\nu u, v \rangle|_{\partial M} = \langle ti_\nu u, tv \rangle$$

After a computation (first assuming that u and v are k -forms and then summing over k), (5.3) can be written as

$$(5.4) \quad (u|\Delta v)_M - (\Delta u|v)_M = (tu|ti_\nu dv)_{\partial M} + (t * u|ti_\nu d * v)_{\partial M} \\ + (t\delta u|ti_\nu v)_{\partial M} + (t\delta * u|ti_\nu * v)_{\partial M}.$$

Now if $u \in H_\Delta$, we wish to define the relative boundary values of u by (5.4). Given any $(f, h) \in H^{\frac{3}{2}}(\partial M, \Lambda \partial M) \times H^{\frac{1}{2}}(\partial M, \Lambda \partial M)$, we use Lemma 5.3 to choose $v = v_{f,h}$ such that $ti_\nu v = f$, $ti_\nu dv = h$, $ti_\nu * v = ti_\nu d * v = 0$ and $\|v\|_{H^2} \lesssim \|f\|_{H^{3/2}} + \|h\|_{H^{1/2}}$. By (5.4) we have for any $u \in H^2(M, \Lambda M)$

$$(tu|h)_{\partial M} = (u|\Delta v_{0,h})_M - (\Delta u|v_{0,h})_M, \\ (t\delta u|f)_{\partial M} = (u|\Delta v_{f,0})_M - (\Delta u|v_{f,0})_M.$$

Since

$$\|tu\|_{H^{-1/2}} = \sup_{\|h\|_{H^{1/2}}=1} |(tu|h)_{\partial M}|, \quad \|t\delta u\|_{H^{-3/2}} = \sup_{\|f\|_{H^{3/2}}=1} |(t\delta u|f)_{\partial M}|,$$

it follows that

$$\|tu\|_{H^{-1/2}} + \|t\delta u\|_{H^{-3/2}} \lesssim \|u\|_{H_\Delta}, \quad u \in H^2(M, \Lambda M).$$

Thus for any $u \in H_\Delta$, we may use the above formulas to define $(tu, t\delta u)$ as an element of $H^{-1/2}(\partial M, \Lambda(\partial M)) \times H^{-3/2}(\partial M, \Lambda(\partial M))$. Applying this argument to $*u$ shows that $(t * u, t\delta * u)$ are well defined for $u \in H_\Delta$ in a similar way. The identity (5.4) then remains true for any $u \in H_\Delta$ and any $w \in H^2(M, \Lambda M)$.

We now claim that $C^\infty(M, \Lambda M)$ is dense in H_Δ . Given this fact, it follows that the maps $u \mapsto (tu, t\delta u)$ and $u \mapsto (t * u, t\delta * u)$ are bounded from H_Δ to $H^{-1/2} \times H^{-3/2}$ and are

uniquely defined by the corresponding maps acting on C^∞ . Finally, (5.1) and (5.2) are true for $u \in C^\infty(M, \Lambda M)$ by the fact that d commutes with the pull-back to the boundary t and thus they extend to $u \in H_\Delta$ by density.

It remains to show that $C^\infty(M, \Lambda M)$ is dense in H_Δ . Suppose that $u \in H_\Delta$. Then $\Delta u \in L^2(M, \Lambda M)$ and $tu, t*u \in H^{-1/2}(\partial M, \Lambda \partial M)$, so for any $\varepsilon > 0$ there exist $v \in \Omega(M)$ and $f, h \in \Omega(\partial M)$ such that

$$(5.5) \quad \|\Delta u - v\|_{L^2} + \|f - tu\|_{H^{-1/2}} + \|h - t*u\|_{H^{-1/2}} < \varepsilon.$$

Now by Lemma 5.2 there exists a smooth form u' such that $\Delta u' = v$, with $tu' = f$ and $t*u' = h$. Then if $w \in H^2(M, \Lambda M)$ satisfies $tw = t*w = 0$, which implies $ti_\nu w = ti_\nu * w = 0$, we get by (5.4) that

$$(u - u' | \Delta w)_M = (\Delta(u - u') | w)_M + (t(u - u') | ti_\nu dw)_{\partial M} + (t*(u - u') | ti_\nu d*w)_{\partial M}.$$

Combining this with inequality (5.5) we obtain

$$|(u - u' | \Delta w)_M| \leq C\varepsilon \|w\|_{H^2}.$$

Now by Lemma 5.2 every form $z \in L^2(M, \Lambda M)$ can be written as Δw for some w with $tw = t*w = 0$. Moreover, we will have

$$\|w\|_{H^2} \lesssim \|z\|_{L^2}.$$

Thus we get that

$$\|u - u'\|_{L^2} \lesssim \varepsilon$$

and hence

$$\|u - u'\|_{H_\Delta} \lesssim \varepsilon.$$

Therefore there is a sequence of smooth forms u_k such that $u_k \rightarrow u$ in H_Δ . \square

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DEPARTMENT OF MATHEMATICS, UNIVERSITY OF KENTUCKY, LEXINGTON, USA

DEPARTMENT OF MATHEMATICS & STATISTICS, UNIVERSITY OF HELSINKI, HELSINKI, FINLAND

DEPARTMENT OF MATHEMATICS & STATISTICS, UNIVERSITY OF JYVÄSKYLÄ, JYVÄSKYLÄ, FINLAND

SCHOOL OF MATHEMATICS AND STATISTICS, SYDNEY UNIVERSITY, SYDNEY, AUSTRALIA