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- Author(s): Sorri, Juha; Greenlees, Paul; Papadakis, Philippos; Konki, Joonas; Cox, Daniel; Auranen, Kalle; Partanen, Jari; Sandzelius, Mikael; Pakarinen, Janne; Rahkila, Panu; Uusitalo, Juha; Herzberg, R.-D.; Smallcombe, J.; Davies, P.J.; Barton, C.J.; Jenkins, D.G.
- Title:Determination of absolute internal conversion coefficients using the SAGE
spectrometer
- Year: 2016

Version:

Please cite the original version:

Sorri, J., Greenlees, P., Papadakis, P., Konki, J., Cox, D., Auranen, K., Partanen, J., Sandzelius, M., Pakarinen, J., Rahkila, P., Uusitalo, J., Herzberg, R.-D., Smallcombe, J., Davies, P.J., Barton, C.J., & Jenkins, D.G. (2016). Determination of absolute internal conversion coefficients using the SAGE spectrometer. Nuclear Instruments and Methods in Physics Research Section A: Accelerators, Spectrometers, Detectors and Associated Equipment, 812, 24-32. https://doi.org/10.1016/j.nima.2015.12.041

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Determination of absolute internal conversion coefficients using the SAGE spectrometer

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Abstract

A non-reference based method to determine internal conversion coefficients using the SAGE spectrometer is carried out for transitions in the nuclei 154 Sm, 152 Sm and 166 Yb. The Normalised-Peak-to-Gamma method is in general an efficient tool to extract internal conversion coefficients. However, in many cases the required well-known reference transitions are not available. The data analysis steps required to determine absolute internal conversion coefficients with the SAGE spectrometer are presented. In addition, several background suppression methods are introduced and an example of how ancillary detectors can be used to select specific reaction products is given. The results obtained for ground-state band E2 transitions show that the absolute internal conversion coefficients can be extracted using the methods described with a reasonable accuracy. In some cases of less intense transitions only an upper limit for the internal conversion coefficient could be given.

Keywords: electron spectroscopy, background subtraction, energy reconstruction, internal conversion coefficient, silicon detector

Preprint submitted to NIM A 19 July 2015

Received in revised form December 7, 2015

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1 1. Introduction

The analysis methods described in this paper have been developed primarily 2 for use with the SAGE (Silicon And GErmanium) spectrometer located at the 3 Accelerator Laboratory of the University of Jyväskylä (JYFL). The spectrometer was constructed with the aim of performing simultaneous in-beam gamma ray and internal conversion electron spectroscopic studies by combining the JUROGAMII germanium array [1] with an electron spectrometer. Detailed discussions of the construction and performance of the SAGE spectrometer can be found in Refs. [2–4]. Further combination of the SAGE spectrometer with the RITU [5, 6] gas-filled recoil separator and the GREAT [7] focal plane spectrom-10 eter allows the use of the recoil-decay tagging technique [8–10]. The primary 11 beam from the JYFL K-130 cyclotron is used to induce nuclear reactions at 12 the target. Depending on the experiment, the reaction products enter RITU 13 and are either dumped or transported to the GREAT spectrometer at the focal 14 plane. The prompt γ -rays emitted in the de-excitation of the populated nu-15 clei are detected with the germanium detectors of the JUROGAMII array and 16 conversion electrons are transported upstream from the target position by an 17 electromagnetic solenoid and detected with a segmented silicon detector. The 18 electrons travel along a helical path following the magnetic field lines in the 19 solenoid, the radius of which is dependent the electron velocity perpendicular 20 to the magnetic field and the magnetic field strength. A high voltage barrier is 21 used to reduce the extremely high flux of δ -electrons produced by interactions 22 of the ion beam with the atomic electrons of the target material. The silicon de-23 tector segments have small average size, which has a drawback at high electron 24 energies. The electron interaction volume grows large and many of the detected 25 electrons deposit energy in more than one segment. An algorithm designed to 26 reduce background generated by scattering between segments is discussed later 27 in the paper. In addition, the fact that the radius of the helical path followed by 28 the electrons is energy dependent provides an opportunity to perform further 29 background filtering. An algorithm exploiting these properties is also presented. 30

31 2. Experiment details

The nuclear structure properties of ¹⁵⁴Sm were recently discussed by Small-32 combe et al. [11]. The main goal of the experiment (here referred to as S06) was 33 to determine internal conversion coefficients from the excited rotational bands of 34 ¹⁵⁴Sm in order to test the hypothesis that the bands have vibrational (β -band) 35 structure. Coulomb excitation was used to populate the excited energy levels in 36 154 Sm by using an enriched target of 154 Sm which was irradiated with a beam of 37 ¹⁶O. Subsequent to the experiment S06, an additional test experiment ST1 was 38 performed. For the ST1 test run the pre-amplifier signals of the outer segments 39 of the SAGE Si-detector were fed through voltage dividers that increased the 40 maximum detectable energy up to 30-40 MeV from the original 2-2.5 MeV. The 41 reasons for this modification and discussion of the results obtained are presented 42 later in this manuscript. The various parameters related to S06 and ST1 exper-43 iments are summarised in table 1. In the aforementioned work by Smallcombe 44 et al., the analysis of the electron-gamma coincidence data and extraction of the 45 internal conversion coefficients relies on the Normalised-Peak-to-Gamma (NPG) 46 method [12]. The NPG method is based on the observation of known transi-47 tions and therefore can not be applied to cases where these reference points are 48 missing. A standalone procedure to determine internal conversion coefficients 49 (ICCs) from in-beam data along with methods to reduce the underlying elec-50 tron background are introduced in this paper. The discussion relies on the same 51 dataset (S06) as that used by Smallcombe *et al.* and the data-analysis param-52 eters such as trigger conditions, event widths etc. are kept as close as possible 53 to those introduced in Ref. [11] in order to enable direct comparison between 54 results. Partial level schemes of studied nuclei are shown in figure 1. 55

56 2.1. Determination of the gamma-ray and electron detection efficiencies

In order to reduce the total counting rate, it is usual that Sn (0.1 mm) and Cu (0.5 mm) absorbers are placed between the target and the germanium detectors of JUROGAMII. In S06 experiment, the absorbers were removed in

Table 1: Summary of various parameters used in the two experiments presented in this work. U_{HV} is the voltage applied to HV barrier and I_{coils} is the current through the SAGE magnetic coils.

	Targ	Target				SAGE	
Rur	Elen	ElementThickness			Energy	\mathbf{U}_{HV}	I_{coils}
			$[mg/cm^2](enrich$	h.)	[MeV]	$\begin{array}{c} \text{gy } U_{HV} \\ \text{V} \end{bmatrix} [\text{kV}] \\ \hline -20 \end{array}$	[A]
S06	154 S	m	1.5 (99%)	¹⁶ O	65	-20	800
ST1	154 S	m	$1.1 \ (98.69\%)$	¹⁶ O	35/65	-20	800
	$^{154}\mathrm{G}$	d	$4.16~(80 {\lesssim} \%)^a$	$^{16}\mathrm{O}$	65	-20	0

^{*a*} Estimated from γ -ray spectrum.

⁶⁰ order to improve the detection of Sm X-rays. The absolute gamma-ray detec-⁶¹ tion efficiency was measured using calibrated ¹³³Ba, ¹⁵²Eu and ²⁰⁷Bi sources. ⁶² The resulting γ -ray detection efficiency curve is presented in figure 2. Note that ⁶³ the curve deviates from the one shown in Ref. [4] as S06 was run without ab-⁶⁴ sorber foils of the JUROGAMII array. For both the SAGE γ -ray and conversion ⁶⁵ electron efficiency curves the data are fitted with a function of the form

$$\epsilon(E) = Exp[\sum_{i=0}^{n} a_i \times ln(\frac{E}{E_0})], \qquad (1)$$

where a_i and E_0 are fitted coefficients and E is the energy of the γ -ray or internal 66 conversion electron. The electron detection efficiency was determined by using 67 calibrated open ¹³³Ba and ²⁰⁷Bi conversion electron sources. The resulting 68 electron detection efficiency curve is shown in figure 3. It can be seen from 69 the figure 3 that it is possible to improve the efficiency for detection of higher 70 energy electrons by using an add-back procedure. Details of this procedure 71 are discussed later in the text. The calibration runs during the experiment were 72 made with no voltage applied to the HV barrier. The effect of the barrier shown 73 in detail in figure 4 is an average behaviour deduced by fitting data obtained 74 from several measurements carried out subsequent to the S06 experiment. As 75 can be seen from figure 4, the effect of the HV barrier is negligible for electron 76 energies over 200 keV, even when the voltage applied to the barrier is -35 kV. 77

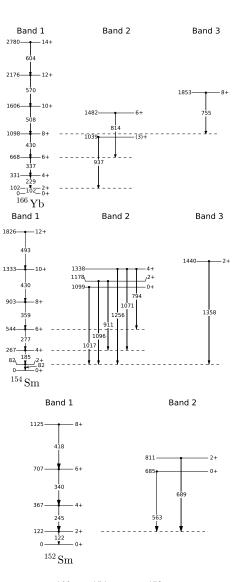


Figure 1: Partial level schemes of ¹⁶⁶Yb, ¹⁵⁴Sm and ¹⁵²Sm. The level and transition energies are rounded to the nearest keV. Data from Ref [13].

78 3. Data analysis procedures

- 79 3.1. Electron add-back/veto
- 80 3.1.1. Description

Around 1 MeV energy the range of the electrons in silicon approaches and

⁸² exceeds the typical segment dimensions of the SAGE silicon detector (1 mm

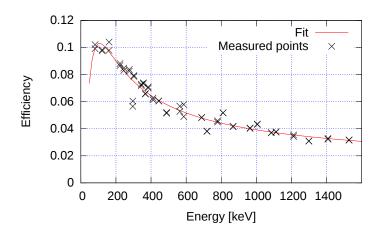


Figure 2: Detection efficiency for single γ -rays in JUROGAM II. The measured points are from calibration runs before and after the experiment. Typical errors on the measured points are $\pm 1\%$ of the value. Error bars have been omitted for clarity.

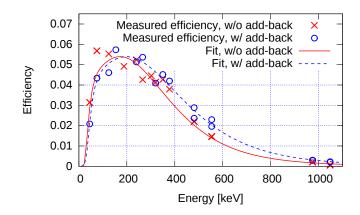


Figure 3: Absolute electron detection efficiency with and without add-back. Typical errors on the measured points are $\pm 3\%$ of the value. Error bars are omitted for clarity.

thick, 1-2 mm wide radially). The calculated 99% stopping range for a 1 MeV
electron in silicon is 1.98 mm. A number of the high energy electrons simply
punch through the 1mm thick detector and deposit only a fraction of their full

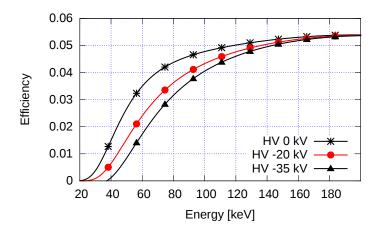


Figure 4: Effect of the HV barrier on the absolute electron detection efficiency. The lowest reliably achieved voltage in experimental conditions is -35 kV.

energy. Moreover, a significant fraction of the high energy electrons scatter to 86 adjacent segments and thus deposit energy in two or more segments. These 87 events can be either recovered by summing the energies (add-back) or can be 88 removed (vetoed) from the data with a simple algorithm. The outline of the add 89 back/veto algorithm is shown in figure 5. An example of a typical spectrum 90 generated by the add-back algorithm is shown in figure 6. In the cases where 91 two or more pixels are hit, the energy deposited in a single pixel is found to 92 have any value up to the maximum for the transition. The inset in figure 6 93 shows the effect of the add-back on the resulting peak shape. Unfortunately, 94 due to the structure of the Si-detector, some electrons lose energy in the inactive 95 area between segments which cannot be detected and a spurious lower energy 96 component is observed in the spectrum. Due to technical issues with bias source 97 the silicon detector bias was limited to 90 V during the measurements and based 98 on the detector I-V curves there is reason to believe that the detector is not 99 completely depleted. This may explain the apparent high energy loss between 100 segments shown as a difference in energy between peaks A and B in figure 6. 101 Note that summing events that first scatter between segments and then escape 102 the detector would yield a continuous tail below the full energy peak and cannot 103

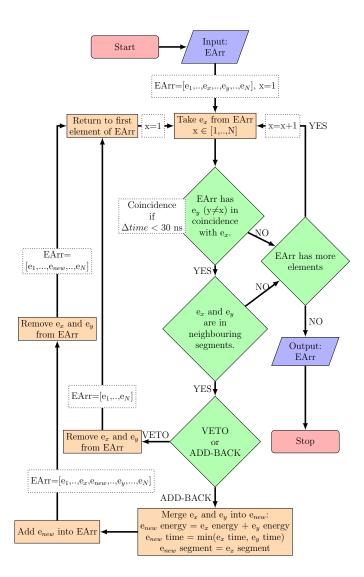
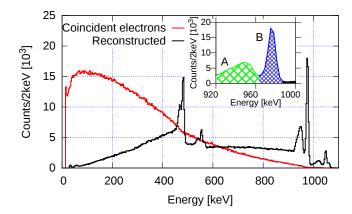


Figure 5: Outline of the add back/veto algorithm. EArr is an iterable data structure where all the elements (electrons) have energy, time and position (segment) attributes.

¹⁰⁴ be the main process generating the secondary peak. Tests conducted during the
¹⁰⁵ commissioning phase showed that the detector does not suffer from cross talk
¹⁰⁶ between signal strip wires or electronics that could explain the secondary peak.
¹⁰⁷ High energy, "punch through" events that effect the common ground behind the
¹⁰⁸ silicon detector can be also ruled out as a source of the secondary peak as the



effect in common ground dissipates over all the silicon segments.

Figure 6: Conversion electron energy spectrum measured with ²⁰⁷Bi source reconstructed by the add-back algorithm along with the original energy spectrum for events where two or more electrons are detected simultaneously in neighbouring segments. *Inset* A typical peak shape in the reconstructed spectrum. Peak A corresponds to electrons that are scattered through the inactive part of the detector which separates the segments. Peak B corresponds to the full energy peak.

109

¹¹⁰ 3.1.2. Neighbouring segments

The SAGE detector is segmented in 90 individual segments and there are 111 several different ways in which it is possible to associate the segments within 112 the add-back procedure. A number of different schemes were tested and the 113 optimal scheme was found to be where all the neighbouring segments within the 114 central region of the detector are grouped together and thereafter those with 115 the longest common borders. The emphasis in determining the best scheme 116 was placed on maximizing the full energy peak areas while at the same time 117 keeping the summing of full energy events with background to a minimum. 118 An outline of the SAGE silicon detector which shows how the different segment 119 types are searched for coincident events is illustrated in figure 7. A comparison 120 of spectra of electrons emitted from ²⁰⁷Bi after application of the add-back 121 and veto algorithms is shown in figure 8. As can be seen from figure 8(a), at 122 energies around 550keV the add-back algorithm has little effect on the efficiency 123

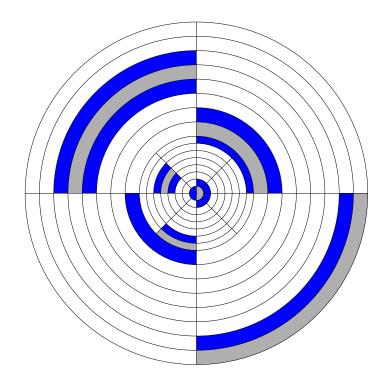


Figure 7: Add-back search patterns for six different segment types. The gray color (lighter color) shows the segment where the event with algorithm index x is detected (see fig. 5), the adjacent blue segments are those which are searched for coincident events.

(electron ranges are shorter) but the use of vetoing significantly reduces the 124 background under the peaks. In figure 8(b) it can be seen that the add-back 125 algorithm increases the efficiency at higher energies, but the effect of vetoing is 126 rather limited. The overall effect of the add-back procedure on the detection 127 efficiency is shown in figure 3. The veto algorithm does not have an effect 128 on the detection efficiency, but can be used to reduce the background in the 129 spectrum. A similar add-back method has been devised for the SPICE electron 130 spectrometer [14]. 131

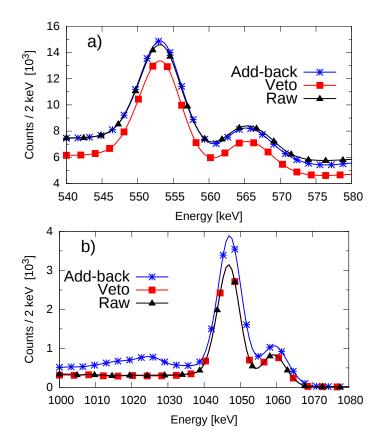


Figure 8: A comparison of the results of applying the veto and add-back algorithms on the conversion electron peaks from ²⁰⁷Bi. a) 554 and 566 keV peaks. b) 1048 and 1060 keV peaks. Data is smoothed to allow better differentiation between the cases.

¹³² 3.2. Filtering with detection radius

As mentioned in the introduction, the operational principle and the design of 133 the SAGE spectrometer gives an opportunity to filter the predominant electron 134 background by studying the detection radius of the electrons as a function of 135 energy. In order to develop the filter, the electron transport properties of SAGE 136 were first probed with standard open electron sources. By using the source data 137 the maximum allowed radius for electrons of a certain energy can be deduced. 138 The source data was used to determine the maximum radius as a function of 139 energy, which was subsequently used to fit a curve based on the function for 140

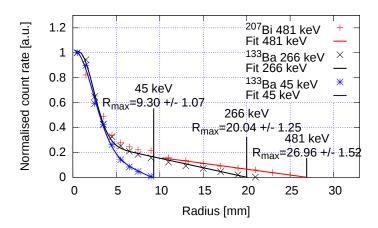


Figure 9: Distributions of electrons with various energies over the SAGE Si-detector. Fits are made according to equation 3 except that only the tail end of 481 keV fit is shown for clarity. Note that the outer radius of the SAGE Si-detector is 24 mm.

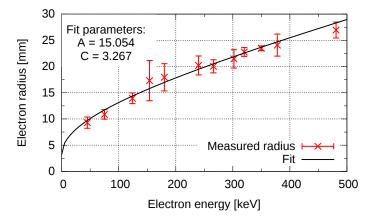


Figure 10: Measured maximum radius as a function of electron energy.

the Larmor radius. Using the relativistic form of the Larmor radius with the assumption that electron velocity is perpendicular to the magnetic field we get

$$R_e(E) = \frac{\beta c \gamma m_e}{eB} = \frac{m_e c \gamma \sqrt{1 - 1/\gamma^2}}{eB}$$

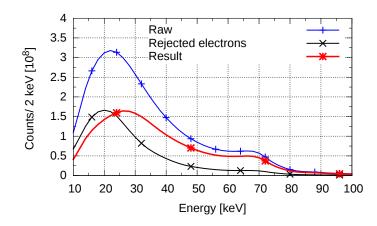


Figure 11: Effect of radial filtering on electron singles spectrum obtained from the S06 inbeam data.

$$= \frac{m_e c \sqrt{\gamma^2 - 1}}{eB} = A \sqrt{(1 + \frac{E}{m_e c^2})^2 - 1} + C, \qquad (2)$$

where m_e is the electron mass, v is the velocity of the electron, c is the speed of 143 light, $\beta = v/c$, $\gamma = 1/\sqrt{1-\beta^2}$, E is the kinetic energy of the electron, e is the elec-144 tron charge and B the magnetic field. The fitting parameters are A $(=m_e c/eB)$ 145 and C. As the direction and the strength of the magnetic field are not clearly 146 defined close to the SAGE silicon detector the last form of equation 2 is used 147 as a basis for fitting. If the electron distribution is well centred electrons with 148 energy E and higher radius than $R_e(E)$ are assumed to have been scattered or 149 generated by beam halo effects and considered to contribute in the background 150 and can be filtered. An example of electron distributions with different energies 151 is shown in figure 9. The fit function used to describe the electron distribution 152 over the radius of the Si-detector has the form 153

$$n(r) = a \times Exp(-b(r - r_0)^2) + cr + d,$$
(3)

where a, b, c, d and r_0 are fitted parameters and r is the radius in mm. To avoid artefacts arising from radial segmentation of the Si-detector the electron radius is randomized within the radial segment limits. For example electron hitting one of the two center segments gets radius within range of [0,1] mm. In an experiment,

the position of the beam spot on the target must be carefully adjusted in order 158 to centre the electron count rate distribution at the detector. Note that the 159 beam spot size and the active spot size in the calibration sources is roughly the 160 same ($\oslash \sim 3$ mm). Originally the magnetic configuration of SAGE caused the 161 electron distribution to systematically veer down and right from the Si-detector 162 center (looking from the target). This was corrected by modifying magnetic 163 shielding around the magnetic coils (see details in Ref's [3, 15]). The δ -electron 164 background is not directly filtered because it is generated in the correct position. 165 The measured maximum radius behaviour determined using the various electron 166 energies from ¹³³Ba and ²⁰⁷Bi sources is shown in figure 10 along with a fit using 167 equation 2. The filter is shown to reduce low energy background below 50 keV 168 by approximately 10% when using source data. The effect of radial filtering on 169 in-beam data is much more prominent as seen in the figure 11. The majority of 170 the background is from δ -electrons produced by interaction of the beam with the 171 target. As in general, the current work focuses on internal conversion coefficients 172 with transition energies higher than 50 keV the effect on the present results is 173 limited. Nevertheless, the filter is employed as it reduces the number of events 174 in the γ -e⁻ coincidence matrix thus easing the analysis. Note that if the add-175 back/veto algorithm is used in the same analysis process with radial filtering the 176 add-back/veto must be performed first in order to avoid errors with the radial 177 filtering arising from scattered events. 178

179 3.3. Definition of coincidence time gates

In order to extract accurate absolute internal conversion coefficients, the 180 time gates for γ - γ and γ - e^- coincidences must be carefully selected. In the 181 present work, the γ - γ and γ - e^- time differences were found to be energy de-182 pendent. The common practice of selecting a single time independent time gate 183 from $\gamma - \gamma$ and $\gamma - e^{-}$ time difference spectra can be lacking in this case. The cor-184 rect time gate can be found by slicing the time spectra in sections and studying 185 the relative number of coincident events within this slice compared to the total 186 number of counts in the coincident peak. As an example, relative peak curves 187

from $^{154}\mathrm{Sm}$ 82-185 keV, 185-277 keV, $^{166}\mathrm{Yb}$ 102-430 keV and random $^{154}\mathrm{Sm}$ 188 $82-^{152}$ Sm 122 keV coincidences are presented in figure 12. The time gate is 189 defined to be the time interval where the relative coincidence peak area is larger 190 than random peak area. As an example the 154 Sm 82-185 keV coincidence gives 191 a high limit of 100ns and a low limit of -60ns (points are circled in figure 13). 192 Due to the experimental timing logic the low limit is related to de-excitation 193 observed first and the high limit to the second in the coincidence cascade in 194 the γ - γ data. Timing of electrons is less affected by electron energy and the 195 gate limits can be set as a function of γ -ray energy with γ - e^- data. Time gates 196 defined from ¹³³Ba, ²⁰⁷Bi, ¹⁵⁴Sm and ¹⁶⁶Yb data are shown in figures 13 and 14. 197 The fits presented in figures 13 and 14 have a general form of 198

$$f(E) = A1 \times Exp(\frac{B1}{\sqrt{E}}) + C1, \qquad (4)$$

where E is the energy in keV and A1, B1, C1 are the fitted parameters.

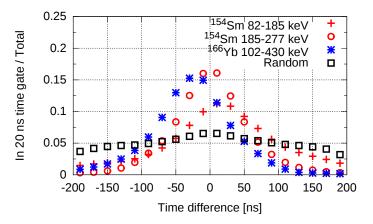


Figure 12: Relative coincident peak size curves compared to total number of coincident counts within [-200:200]ns time gate.

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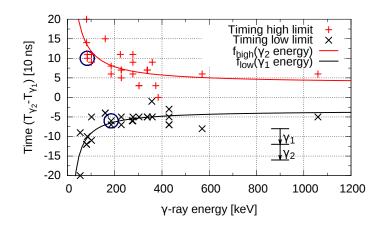


Figure 13: Energy-dependent time gate limits for γ - γ coincidences.

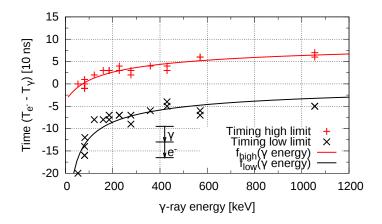


Figure 14: Energy-dependent time gate limits for γ - e^- coincidences. Both limit fits are functions of γ -ray energy.

200 4. Results

201 4.1. Extraction of internal conversion coefficients

202 In this work the internal conversion coefficient α_{exp} is determined from

$$\alpha_{exp} = \frac{N_e \times \epsilon_{\gamma}}{N_{\gamma} \times \epsilon_e},\tag{5}$$

where N_e is the number of detected electrons, N_{γ} is the number of detected γ -rays, ϵ_{γ} and ϵ_e are the detection efficiencies for γ -rays and electrons, respectively. In this work, the effect of angular correlations is neglected. Using the

data obtained in the S06 experiment, a large number of experimental internal 206 conversion coefficients could be determined. The measured absolute internal 207 conversion coefficients for ¹⁵⁴Sm, ¹⁵²Sm and ¹⁶⁶Yb ground state band transi-208 tions that are all considered to be pure E2 character as a function of electron 209 energy are shown in figure 15. The results derived from the raw (no algorithms 210 applied and unfiltered) matrices differ from the reference tabulated values ob-211 tained using BrIcc [16] conversion coefficient calculator. It can be seen that no 212 single normalisation constant would yield agreement throughout the full energy 213 range and without a common factor the NPG method cannot be used. After 214 application of time gates, either add-back or veto and radial filtering for the 215 electron events in the silicon detector the overall result is much more agreeable. 216 A more detailed plot of the final result is shown in figure 16. As can be seen in 217 figure 16 the measured ICCs below 200 keV are systematically lower than the 218 tabulated values. The difference is thought to be the result of the interactions 219 of the electrons with the thick target. The values deduced indicate that ¹⁶⁶Yb 220 is less affected. The difference can arise from the fact that ¹⁶⁶Yb is produced 221 in a fusion-evaporation reaction, meaning that ¹⁶⁶Yb has a kinetic energy of 222 only ~ 6 MeV and range in samarium of ~ 0.9 mg/cm² with beam (¹⁶O) energy 223 of 65 MeV. This should be compared with that for excited 154 Sm which is ~ 4 224 mg/cm^2 . As the beam particles pass through the target matter they lose energy 225 and therefore creation of sub-barrier fusion products deeper in the target mat-226 ter is less likely. As the electrons emitted from ¹⁶⁶Yb travel through less target 221 matter the energy loss is smaller and probability of scattering is lower hence 228 it is more likely that the emitted electrons contribute to the full energy peaks. 229 According to the rule-of-thumb given in Ref [17] the optimal target thickness 230 for measurements of conversion electrons in an energy range of 100 to 500 keV 231 would be $0.3-0.7 \text{ mg/cm}^2$. This is significantly less than the 1.5 mg/cm^2 target 232 used in this case and negative effects on the spectrum quality can be expected. 233 The relative effect of the target on electron transmission is shown figure 17. If 234 we process the ICCs measured below 200 keV with the rough assumption that 235 samarium conversion electrons originate evenly throughout the target depth 236

and ytterbium conversion electrons just from the first half of the target we get
attenuation factors shown in table 2. Applying these estimated target attenuation values to measured ICCs yield values close to those calculated by BrIcc.
In ¹⁶⁶Yb the analysis is further complicated by the overlap of LMN conversion
lines and K conversion lines from different transitions. The standard timing

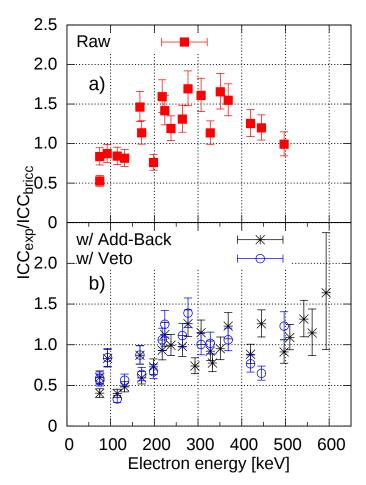


Figure 15: Experimental internal conversion coefficients relative to tabulated values extracted from the data obtained in the S06 experiment. a) Result from raw matrices. b) Results after time gating, electron add-back or veto and radial filtering.

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method where only time independent γ - γ and γ - e^- time gates are used results in a γ - γ gate of [0,60]ns and a γ - e^- gate of [-100,60]ns [11]. Compared to these

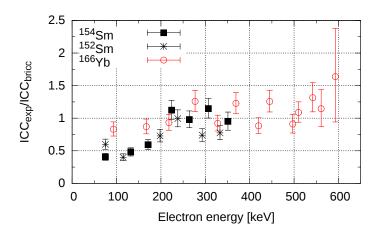


Figure 16: Comparison of the experimental ICCs relative to tabulated values for isotopes of Samarium and ¹⁶⁶Yb obtained with add-back algorithm. The better agreement for ¹⁶⁶Yb below 200 keV can be understood in terms of the reaction kinematics and the interactions of electrons with the target.

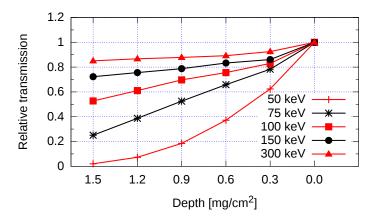


Figure 17: Effect of Sm target on electron transmission. Simulation result taking account of energy loss, scattering and SAGE electron acceptance.

the energy dependent gates shown in figures 13 and 14 are significantly different. If we consider the energy range between 200-400 keV the γ - γ time independent gate is roughly half the width of the energy dependent time gate. According to equation 5 this would yield ICCs approximately a factor of 2 too large as N $_{\gamma}$ is halved. This would explain why the measured ICCs obtained in Ref [11] were

Origin	e^- energy	$\alpha_{Rel}{}^a$	Attenuation	Corrected	
	$[\mathrm{keV}]$		$\operatorname{coefficient}^b$	α_{Rel}	
$^{152}\mathrm{Sm}$	75	0.59(7)	0.6(1)	1.0(2)	
	115	0.39(5)	0.7(1)	0.6(4)	
	198	0.73(9)	0.8(1)	0.9(2)	
154Sm	75	0.41(5)	0.6(1)	0.7(3)	
	132	0.48(6)	0.7(1)	0.7(3)	
	171	0.59(7)	0.8(1)	0.7(3)	
¹⁶⁶ Yb	93	0.83(9)	0.8(1)	1.0(2)	
	167	0.87(9)	0.9(1)	1.0(2)	

Table 2: Approximate attenuation coefficients for ICCs below 200 keV and resulting relative ICCs with results obtained with add-back algorithm.

^a $\alpha_{Rel} = \alpha_{exp} / \alpha_{BrIcc}$

 b Approximated from figure 17

 $_{249}$ 1.8 times the literature value².

250 4.2. Internal conversion coefficients of high-energy transitions

The main goal of the S06¹⁵⁴Sm Coulomb excitation experiment was to study 251 inter-band transitions between the excited side bands and the ground state band. 252 The results for these higher energy transitions are shown in table 3. In several 253 cases only an upper limit could be given due to the lack of statistics. The 254 partial level schemes showing the transitions investigated are shown in figure 1. 255 Data for level schemes are from Ref [13]. The measured ICC for the 4^+_3 to 6^+_1 256 transition in ¹⁵⁴Sm suggests E2 character. However, a M1+E2 transition with 257 a mixing ratio $\delta \leq 0.5$ is possible within the error limits. 258

²J. Smallcombe, private communication

Table 3: Experimental results for α_K ICCs in ¹⁵⁴Sm, obtained by using energy dependant time gates, either the veto or add-back algorithm and filtering. The level and multipolarity assignments are as listed in NNDC [13] if not otherwise stated. The upper limits are deduced with a 90% confidence limit according to Ref [18]. The measured value of the mixing ratio (δ) is given where available.

Origin	γ energy	e^- energy	I_i^{π}	I_f^{π}	σL	δ	$\alpha_{K,exp}$	$\alpha_{K,exp}$	$\alpha_{K,lit}$
	[keV]	$[\mathrm{keV}]$					add-back	veto	
152 Sm	563	516	0_{2}^{+}	2_{1}^{+}	E2	-	0.006(4)	0.005(3)	$0.0069(34)^a$
	689	642	2^{+}_{2}	2_{1}^{+}	E0+M1+E2	8^{+6d}_{-3}	0.003(2)	0.04(2)	$0.0297(75)^a$
154 Sm	795	748	4_{3}^{+}	6_{1}^{+}	$(E2)^{b}$	-	0.0030(27)	$\leq 0.021(1)$	$0.00345(5)^c$
	911	864	2^{+}_{3}	4_{1}^{+}	E2	-	0.0023(20)	0.003(2)	$0.0034(16)^a$
	1017	970	0_{3}^{+}	2_{1}^{+}	E2	-	$\leq 0.038(1)$	$\leq 0.018(1)$	$0.00204(3)^c$
	1071	1024	4_{3}^{+}	4_{1}^{+}	M1+E2	$> 50^{e}$	$\leq 0.025(1)$	$\leq 0.038(1)$	$0.0079^{+0.0087\ a}_{-0.0073}\ ^a$
	1096	1050	2^{+}_{3}	2_{1}^{+}	M1+E2	$30(21)^{f}$	$\leq 0.036(1)$	$\leq 0.043(1)$	$\leq 0.0067(6)^{a}$
	1256	1209	4_{3}^{+}	2_{1}^{+}	E2	-	$\leq 0.061(1)$	$\leq 0.051(2)$	$0.001329(19)^c$
	1358	1311	2_{4}^{+}	2_{1}^{+}	[M1+E2]	$19(10)^{f}$	$\leq 0.036(1)$	$\leq 0.032(1)$	$0.0014(3)^c$
¹⁶⁶ Yb	755	694	8^+_3	8_{1}^{+}	E0+M1+E2	-	0.03(2)	0.005(4)	$0.0158(45)^{\ a}$
	814	753	6_{2}^{+}	6_{1}^{+}	M1	-	0.008(3)	0.008(3)	$0.0069(28)^{\ a}$
	937	876	$(3)_2^+$	2_{1}^{+}	${ m E2}$	-	0.014(6)	$\leq 0.49(2)$	$0.00351(5)^c$

^a Experimental result from Ref [11].

^b Not available in NNDC, own assignment based on the experimental ICC value.

^c BrIcc result. ^d From Ref [19]. ^e From Ref [13]. ^f From Ref [20].

259 4.3. Observation of high-energy events in the silicon detector

As can be seen in figure 18, the electron spectrum measured from the $^{16}O^{+154}Sm$ Coulomb excitation reaction shows a significant number of events that lie outside the dynamic range of the analogue to digital converters in the data acquisition system. The number of these events far exceeds that which is expected based on the behaviour and shape of the electron spectrum at high energies. In a test experiment (ST1), these "overflow" events were shown to be due to the detection of backscattered ¹⁶O beam in the SAGE silicon detector.

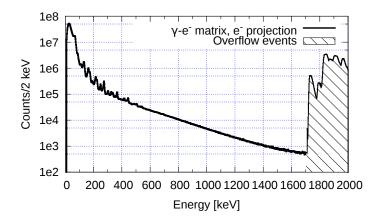


Figure 18: A projection of the electron spectrum from the γ -electron coincidence matrix produced in the ¹⁶O+¹⁵⁴Sm Coulomb excitation reaction. Note the abundance of "overflow" events observed at high energies.

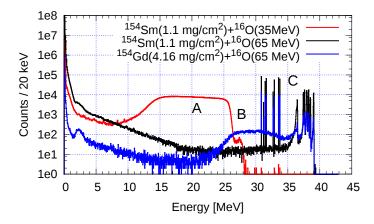


Figure 19: Spectra from the SAGE silicon detector with increased dynamic range from the ST1 test run with a ¹⁵⁴Sm target bombarded by a beam of ¹⁶O at an energy of 35 MeV and 65 MeV and ¹⁵⁴Gd target bombarded with 65 MeV energy. (A) Backscattered ¹⁶O from samarium. (B) Backscattered ¹⁶O from gadolinium. (C) Overflow events.

In the ST1 test experiment, the gain in amplification of the outer segment (59 to 90) signals of the SAGE silicon detector were reduced by voltage dividers allowing increased dynamic range and detection of backscattered ¹⁶O beam particles. An alpha source was used for energy calibration confirming that the maximum energy range had increased to \sim 30-40 MeV from the orig-

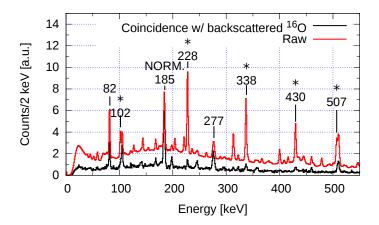


Figure 20: Gamma-projections from γ -electron matrices. Normalized with the area of 185 keV peak. From ST1 test run with 65 MeV beam energy. Peaks marked with asterisk originate from ¹⁶⁶Yb.

inal $\sim 2-2.5$ MeV. The original Coulomb excitation experiment was repeated 272 with beam energies of 35 and 65 MeV. For validation purposes scattering from 273 4.16 mg/cm^{2} ¹⁵⁴Gd target with 65 MeV beam energy was also studied. In the 274 resulting spectrum shown in figure 19 backscattered ¹⁶O is clearly observable. 275 The energy distribution of the scattered particles correspond with calculated 276 distribution from a 1.1 mg/cm² ¹⁵⁴Sm target. The identification of backscat-277 tered ¹⁶O was further confirmed when the beam energy was changed, resulting 278 in a corresponding shift in the energy of the backscattered particles. Selecting 279 events in coincidence with backscattered ¹⁶O, the events arising from sub-barrier 280 fusion such as 166 Yb can be removed. A normalized γ -ray projection coincident 281 with backscattered ¹⁶O ions is shown in figure 20. After demanding the coin-282 cidence, the contribution from 166 Yb (102, 228, 338, 430 and 507 keV peaks) 283 is significantly smaller compared to that from 154 Sm (82, 185 and 277 keV). 284 However, the peak areas are significantly reduced. In the raw projection of the 285 γ -e⁻ matrix the peak area of the 185 keV transition is in the order of 10⁸ but 286 in the gated projection only on the order of 10^3 . Since the statistics obtained in 287 the short test run were rather low, the data were not analysed further. However, 288 with a longer run and improved detection system demanding a coincidence with 289

scattered particles could produce exceptionally clean data for the extraction ofICCs.

²⁹² 5. Conclusions

Experimental internal conversion coefficients have been successfully extracted 293 with non-reference based methods. The crucial step in order to determine ab-294 solute ICC values with reasonable accuracy is the creation of energy-dependent 295 The result obtained through demanding coincidence with backscattime gates. 296 tered ions (Fig 20) resembles greatly the results obtained with recoil gating or 297 recoil-decay tagging in γ -ray spectroscopic studies (see for example Ref [21]). In 298 order to fully exploit the possibilities of this method plans to instrument SAGE 299 with an additional heavy ion detector have been made. One notable detector 300 based on recent developments with optical fibres such as presented in Ref [22] 301 is under consideration. 302

303 6. Acknowledgements

This work has been supported through the UK Science and Technology Facilities Council, the Academy of Finland under the Finnish Centre of Excellence Programme 2006-2011 (Nuclear and Accelerator Based Physics Contract No. 213503), and the European Research Council under the SHESTRUCT project (Grant Agreement No. 203481). The support from GAMMAPOOL network is acknowledged.

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